Towards the six-dimensional view of the Orion Complex

Marina Kounkel,

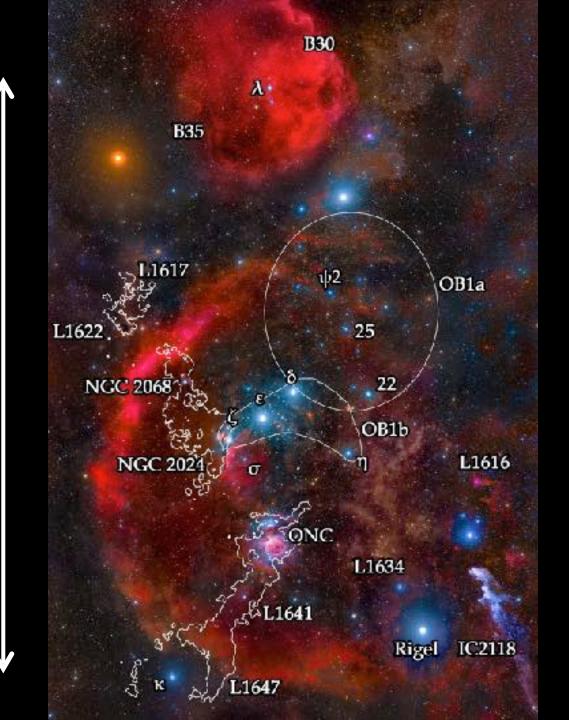
APOGEE

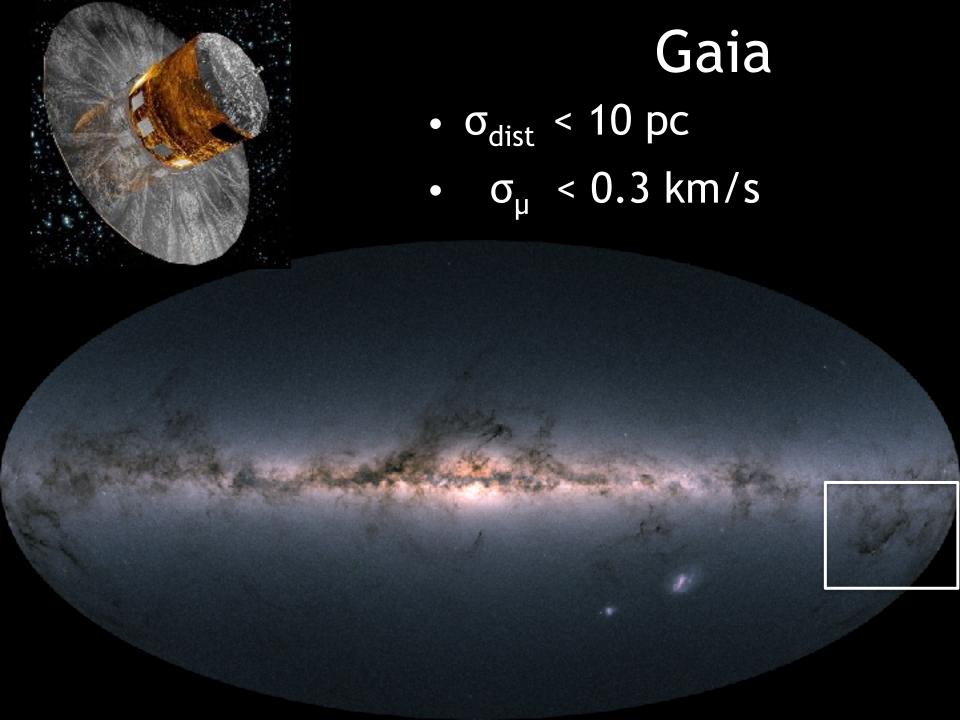
Collaboration

Orion Molecular Cloud Complex

20° ~150 pc

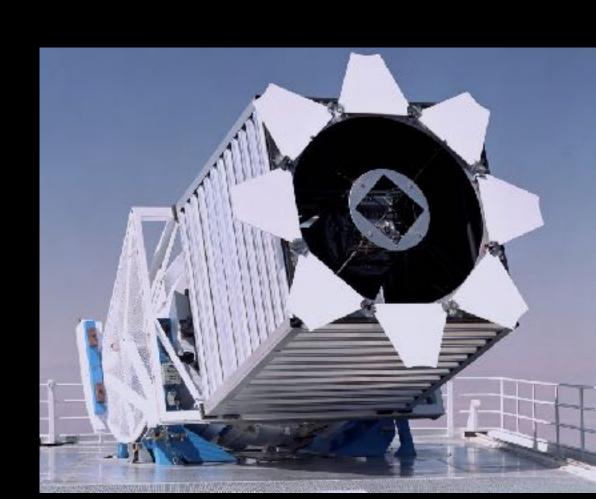
- Multiple stellar populations
- Age range from < 1 to >10 Myr





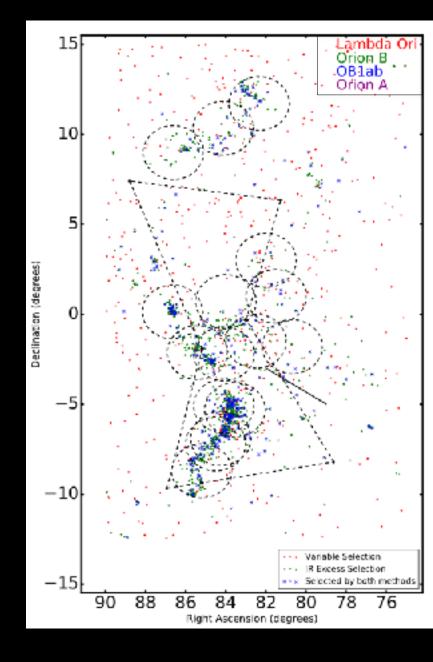
APOGEE

- SDSS 2.5 m telescope
- Multi-object spectrograph
- 3 degree FOV
- 300 fibers
- H-band
- R ~22,500

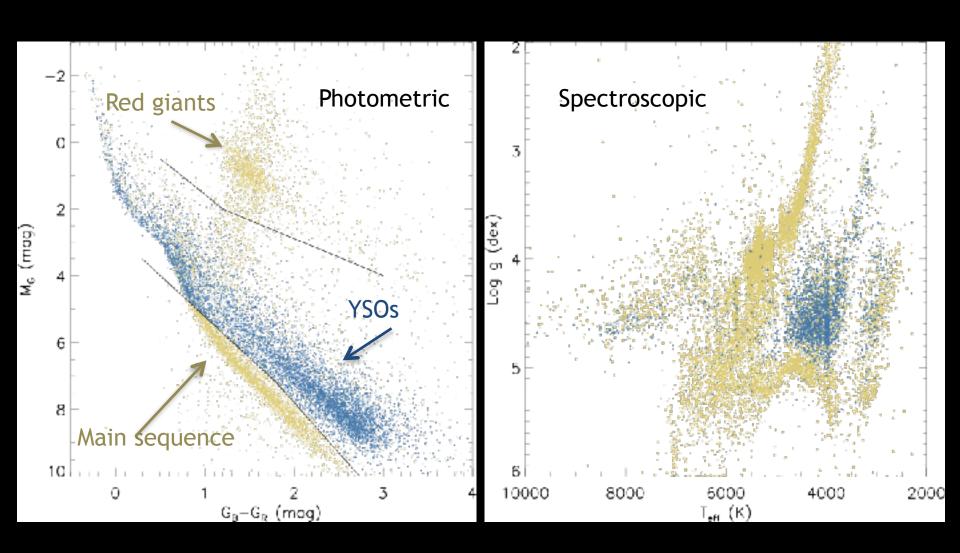


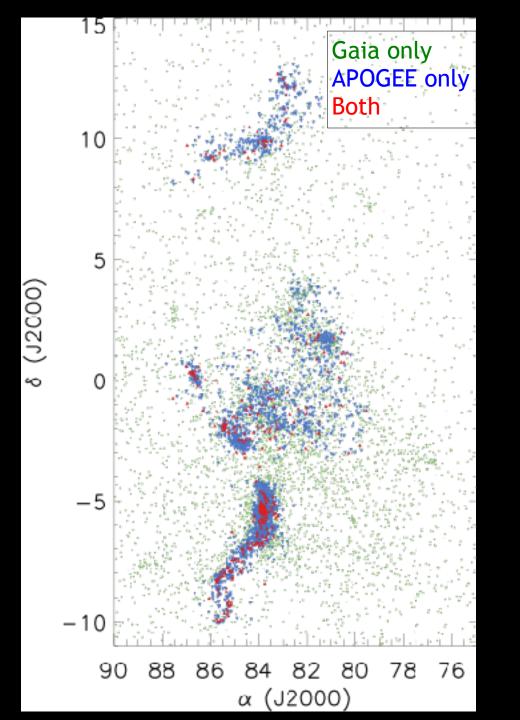
APOGEE Young Cluster Survey - Orion

- ~23,000 spectra
- ~9000 sources
- Uniform target selection
 - IR excess
 - Variability
 - Previous YSO identification

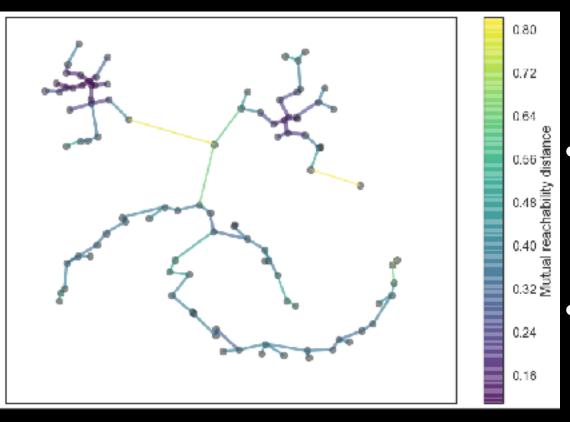


Color cuts



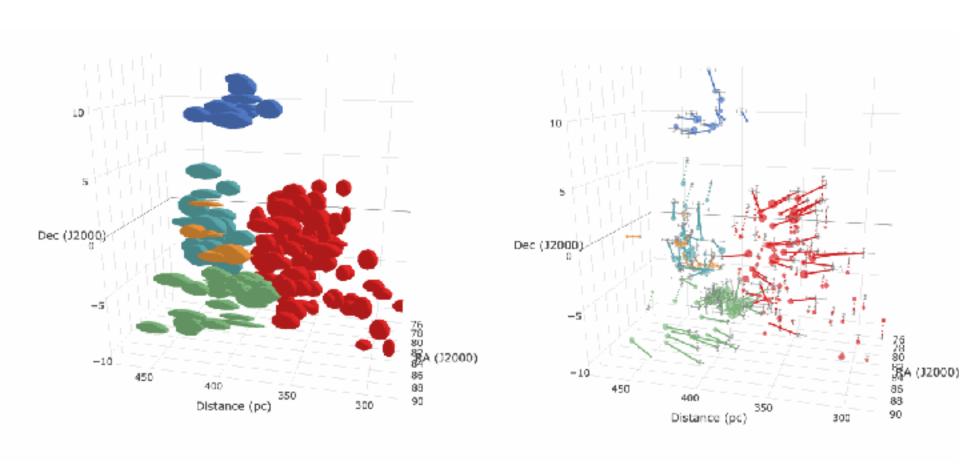


Hierarchical clustering



- Measure distance between all points
- Construct a minimum spanning tree
 - Determine the appropriate place to cut the branches

6D structure



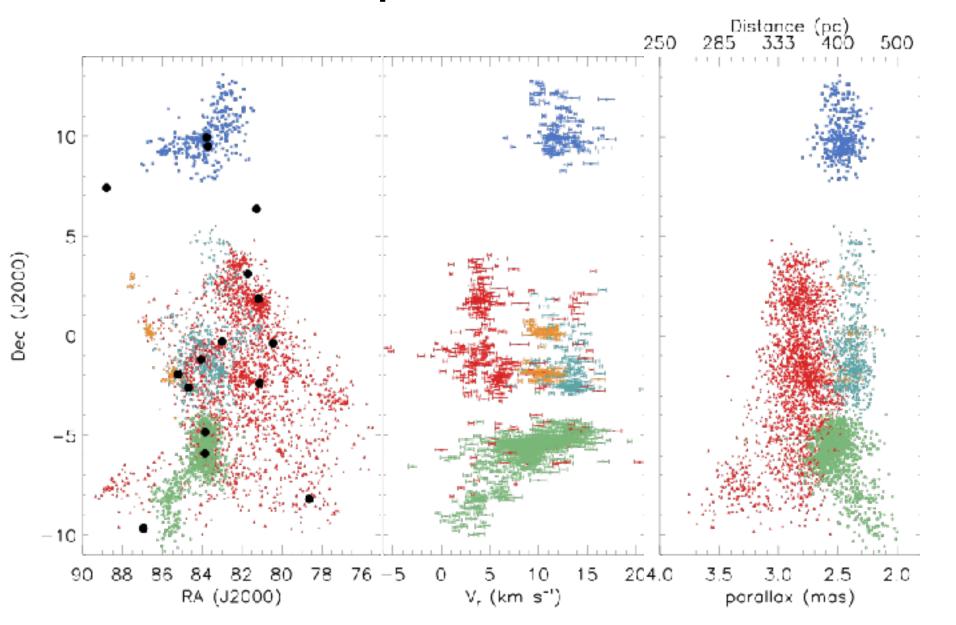
Orion A

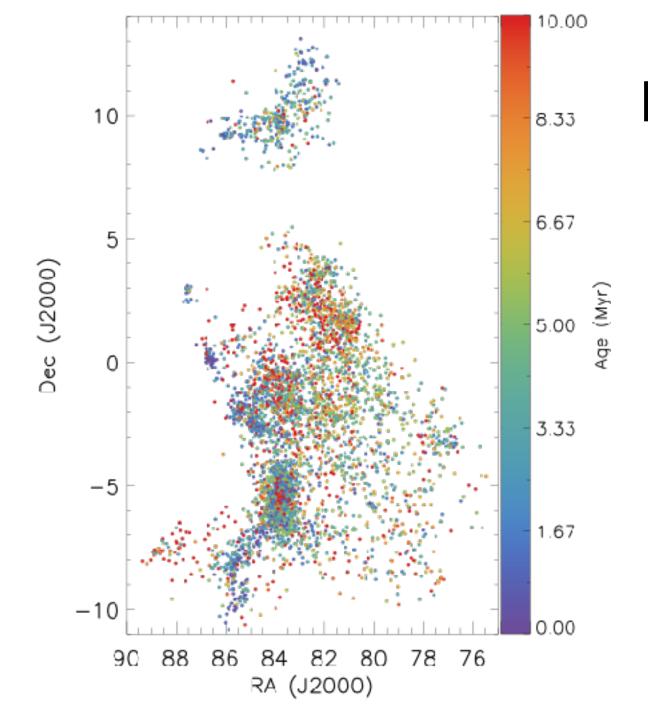
Orion B

Orion C

Orion D

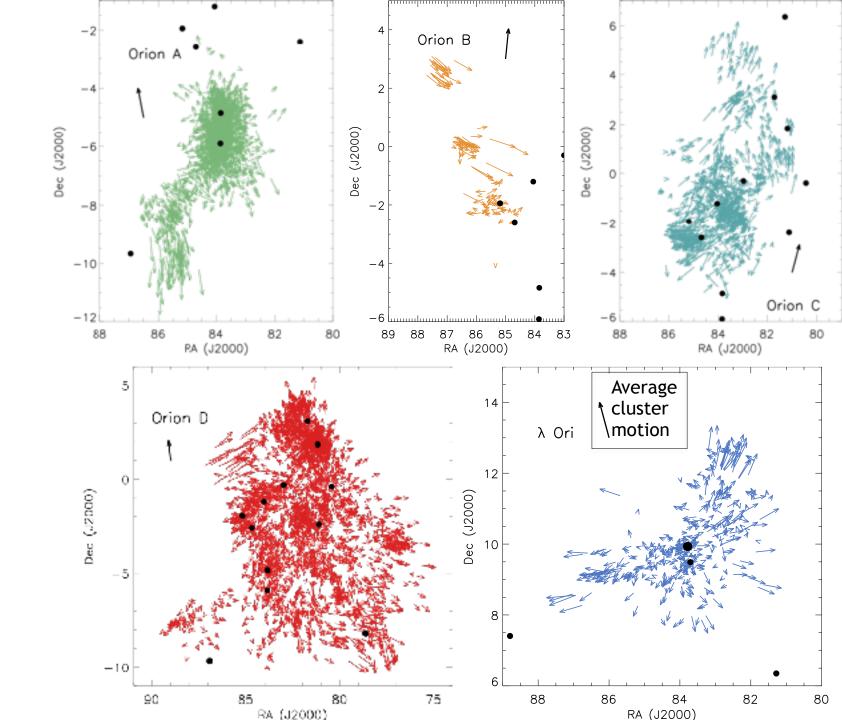
Orion Complex





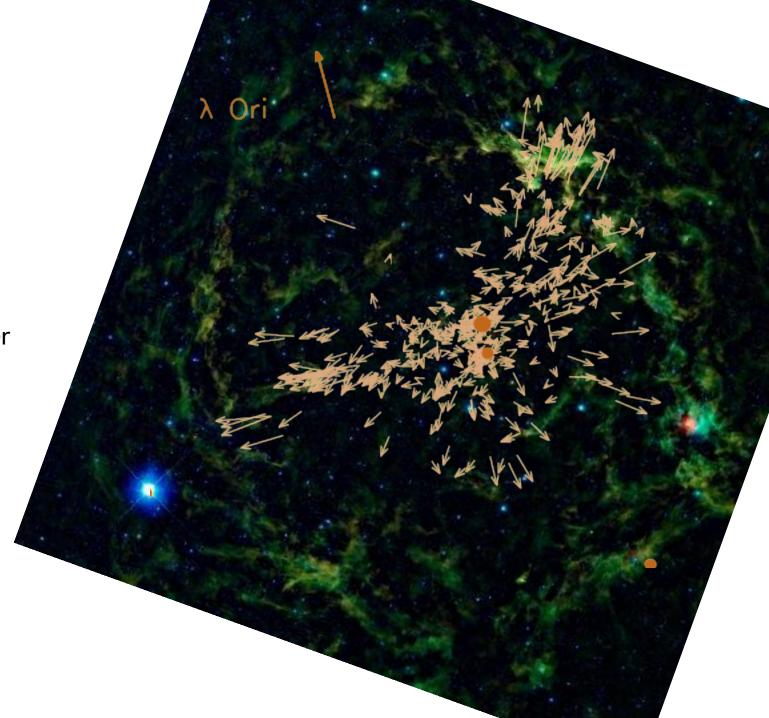
Distribution of ages

Proper Motions



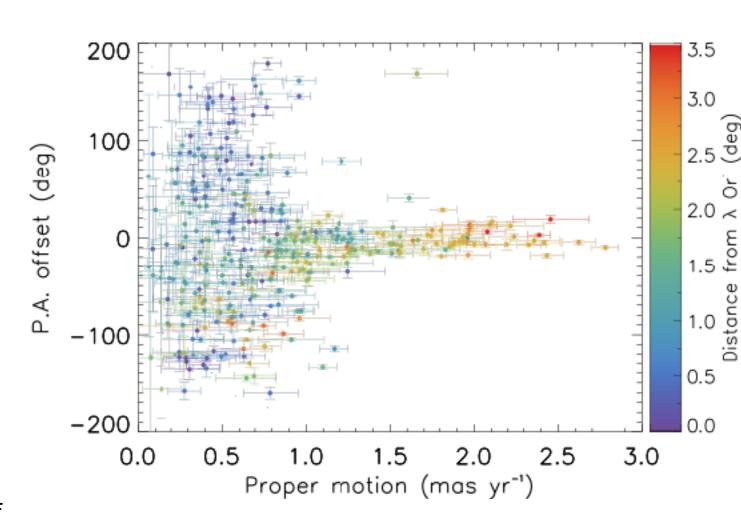
λOri

- Supernova occurred in λ Ori
- Central cluster age ~5 Myr
- Ages near the edge ~ 2 Myr

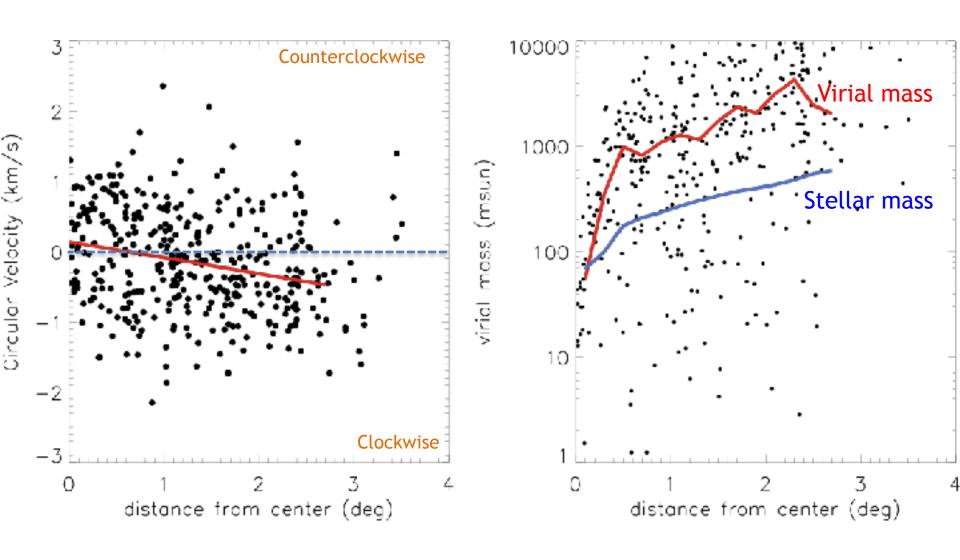


λ Ori

- Supernova occurred in λ Ori
- Central cluster age ~5 Myr
- Ages near the edge ~ 2 Myr
- Stars within 1.5 deg are virialized
- Further than that they are moving radially
- Expansion age of 4.8 Myr



Circular velocity component



APOGEE Collaboration at CS20

- Kevin Covey (Th 16:50, star cluster splinter)
 - Accretion, ages, and multiplicity in the first
 125 Myrs: A systematic view from the
 APOGEE-2 Young Cluster Survey
- Genaro Suárez (Poster 292)
 - System IMF of the 25 Orionis Stellar Group
- Serena Kim (Poster 160)
 - Probing Effect of External UV Radiation on Young Stellar and Substellar Mass Objects

