

Probing the surfaces of Sun-like stars using transiting planets and 3D MHD simulations

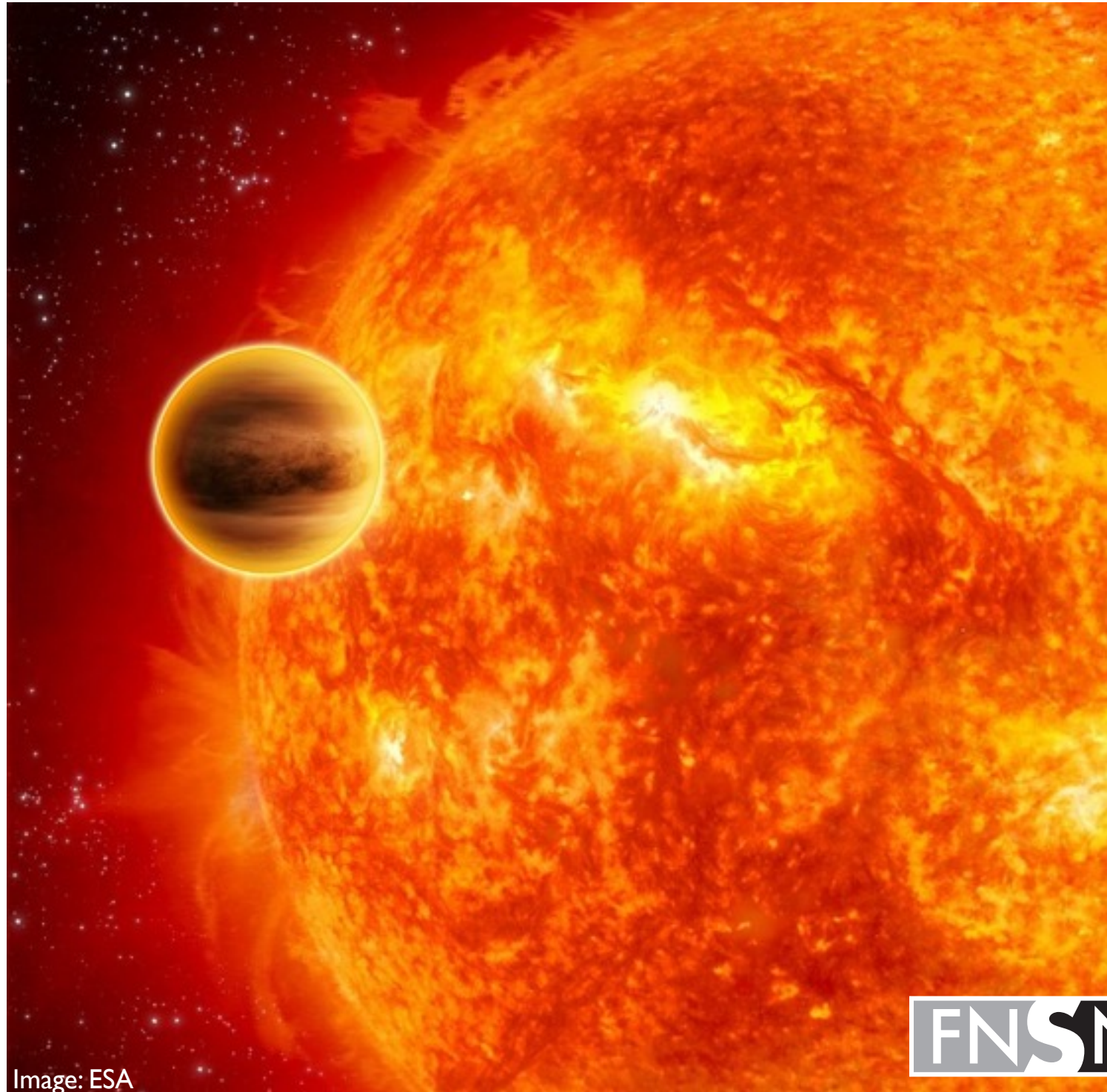


Image: ESA

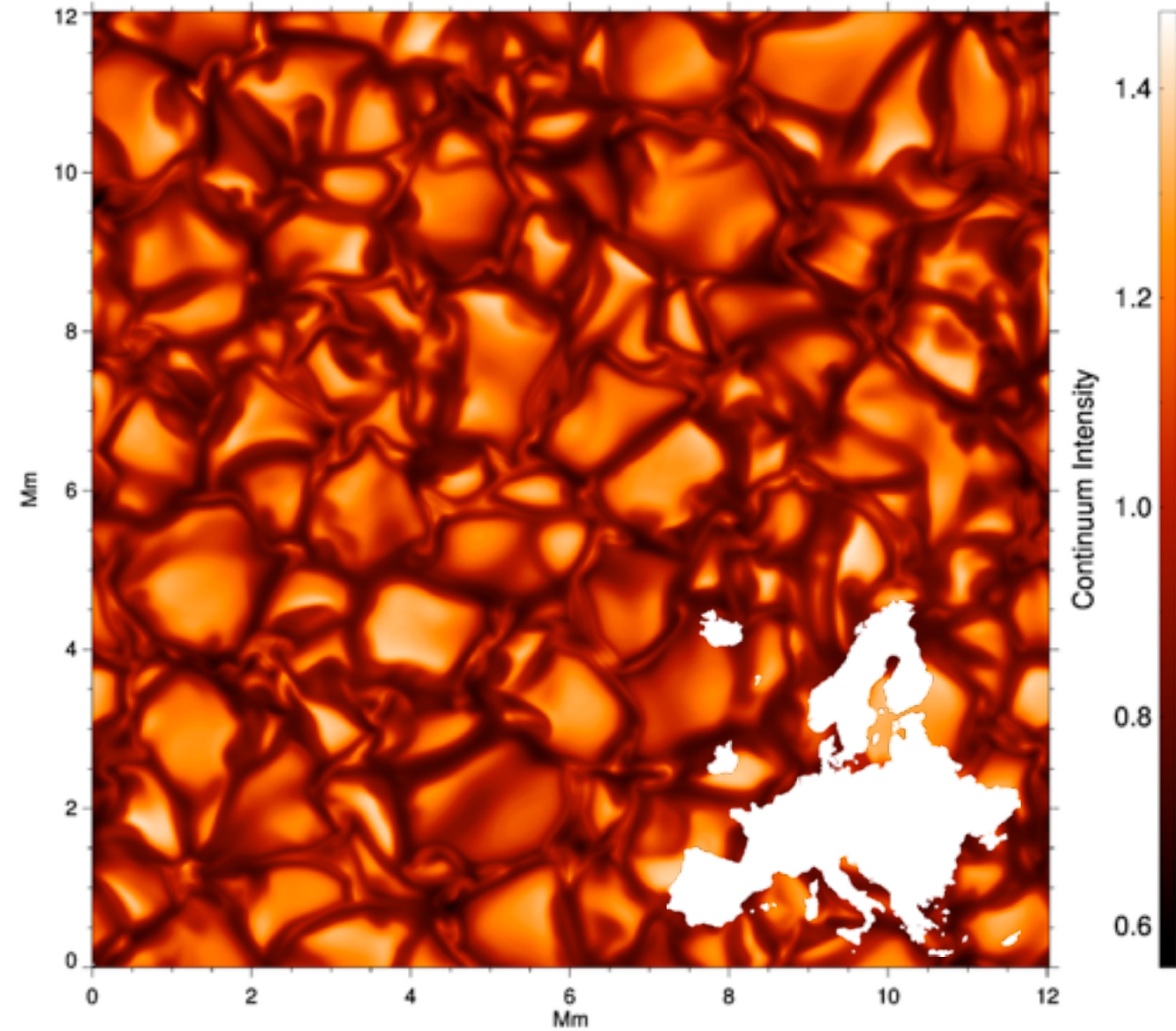
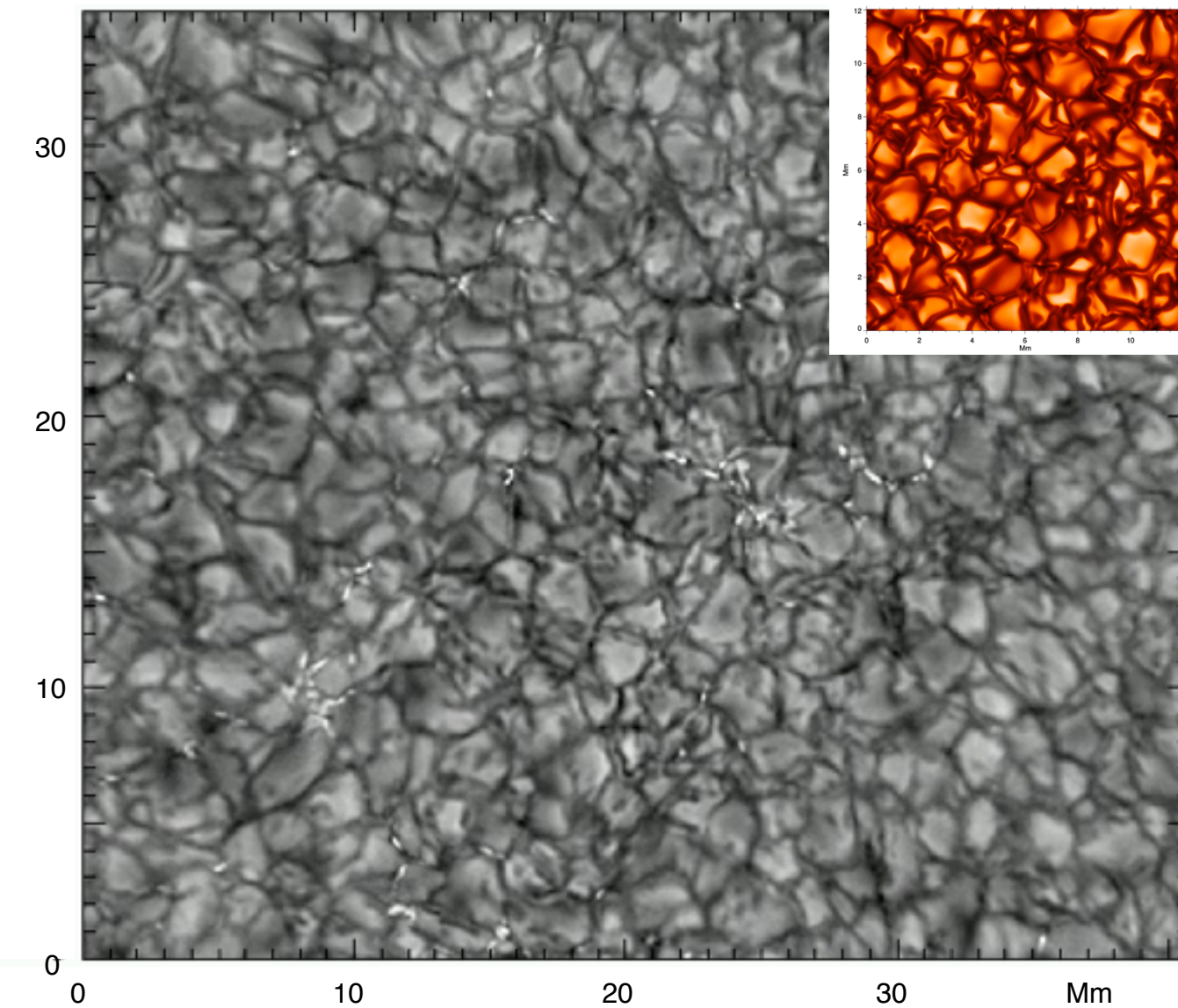
Dr Heather Cegla
CHEOPS Fellow



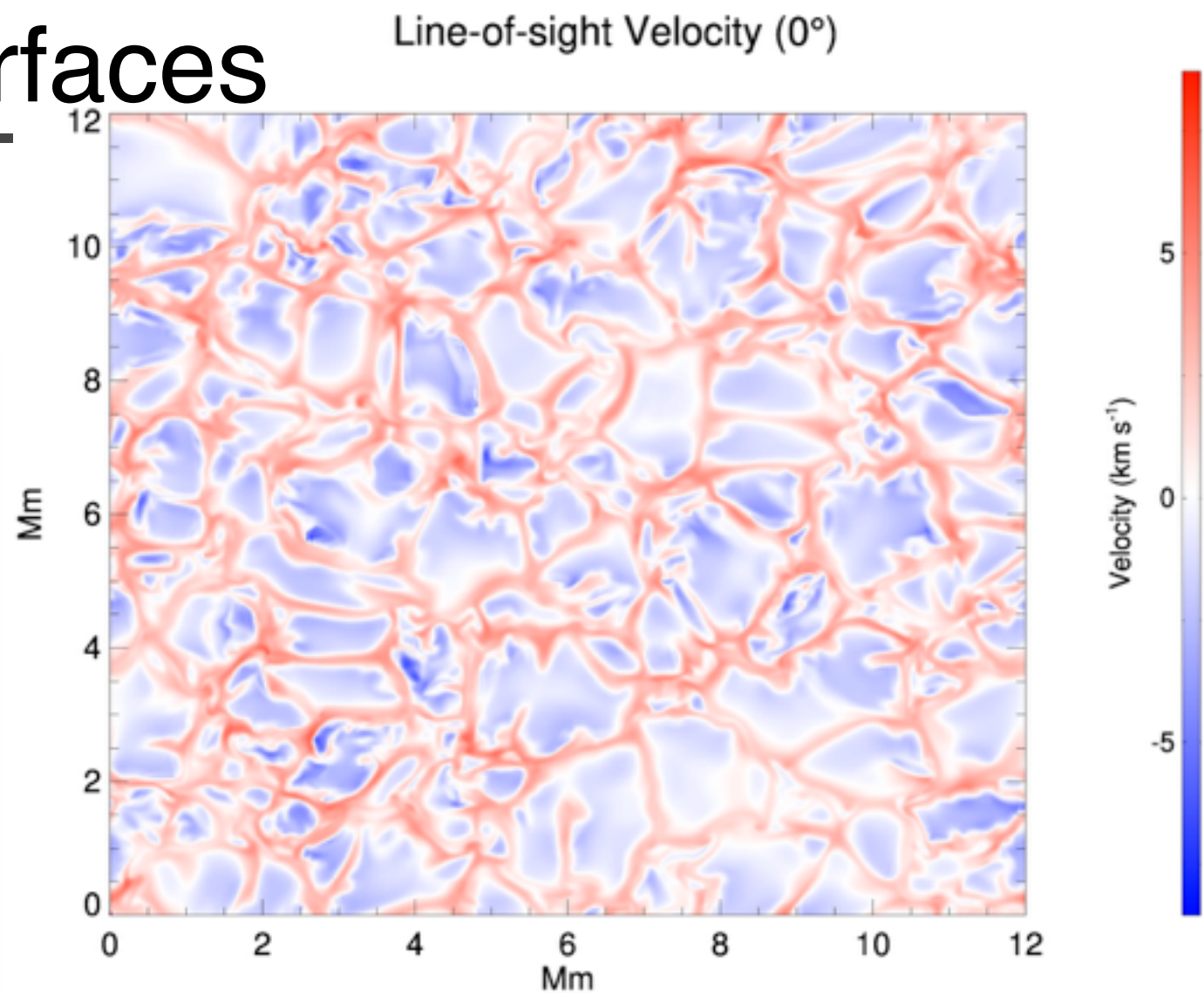
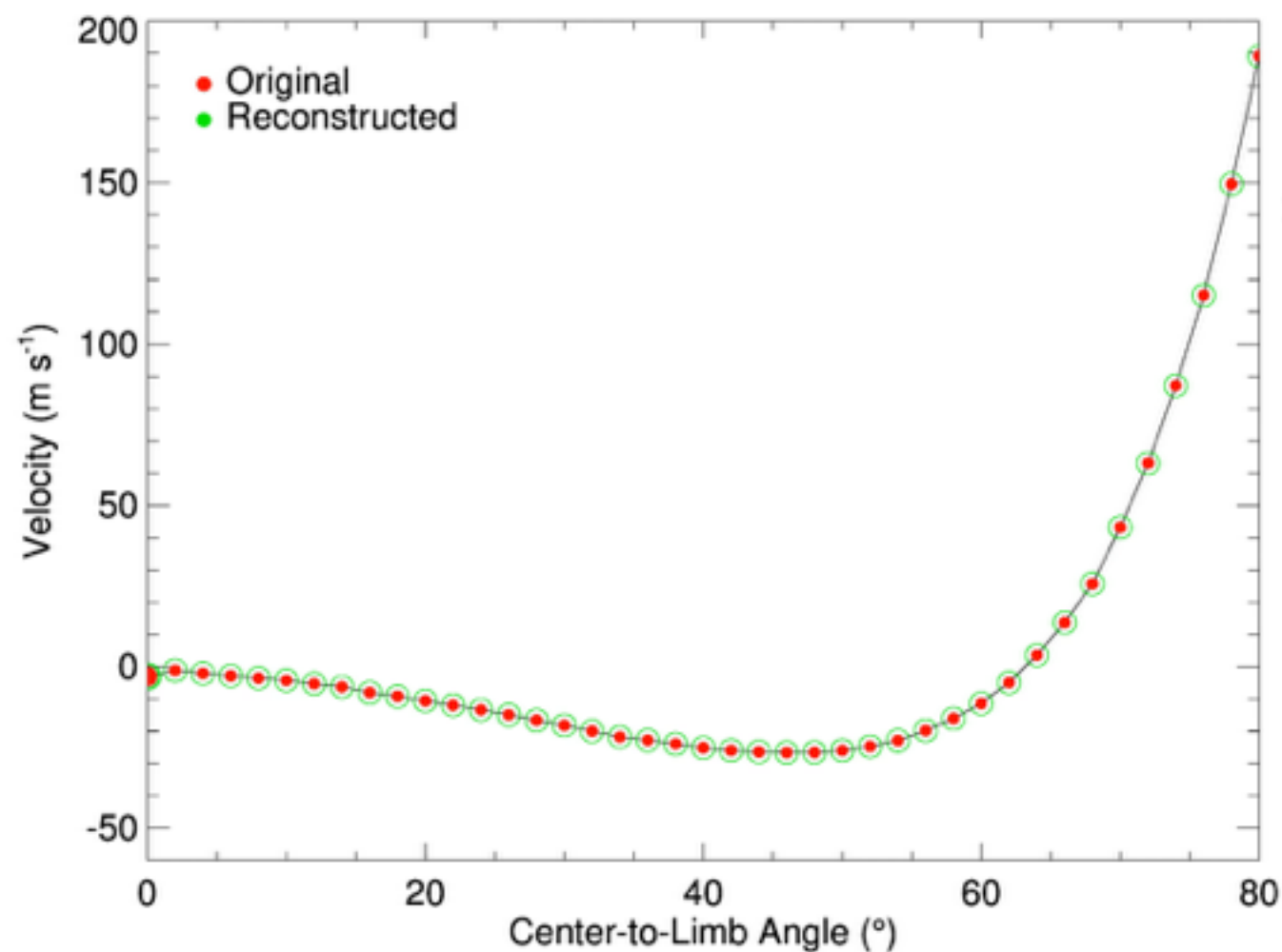
UNIVERSITÉ
DE GENÈVE



The solar surface

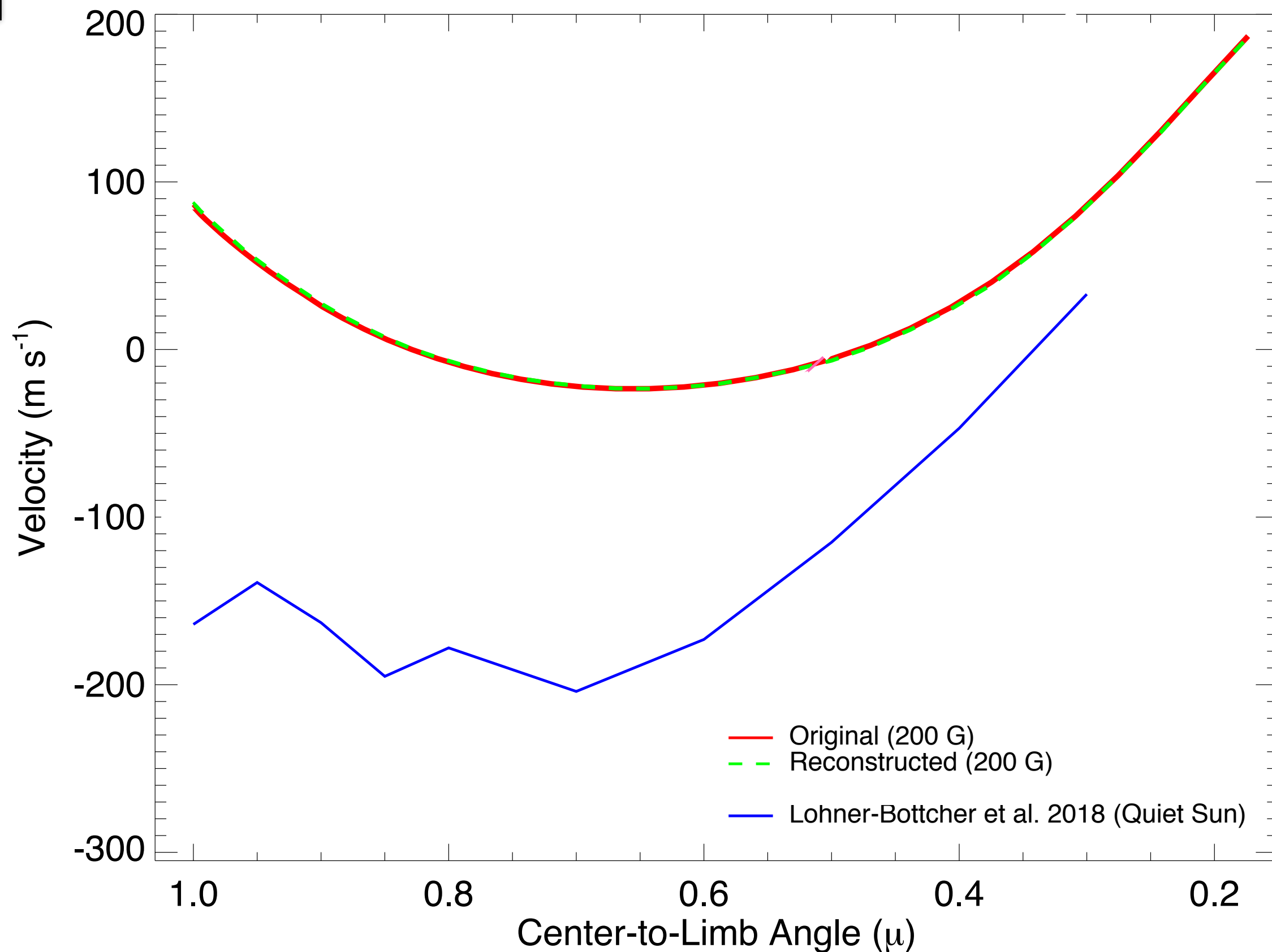


Transiting Stellar Surfaces



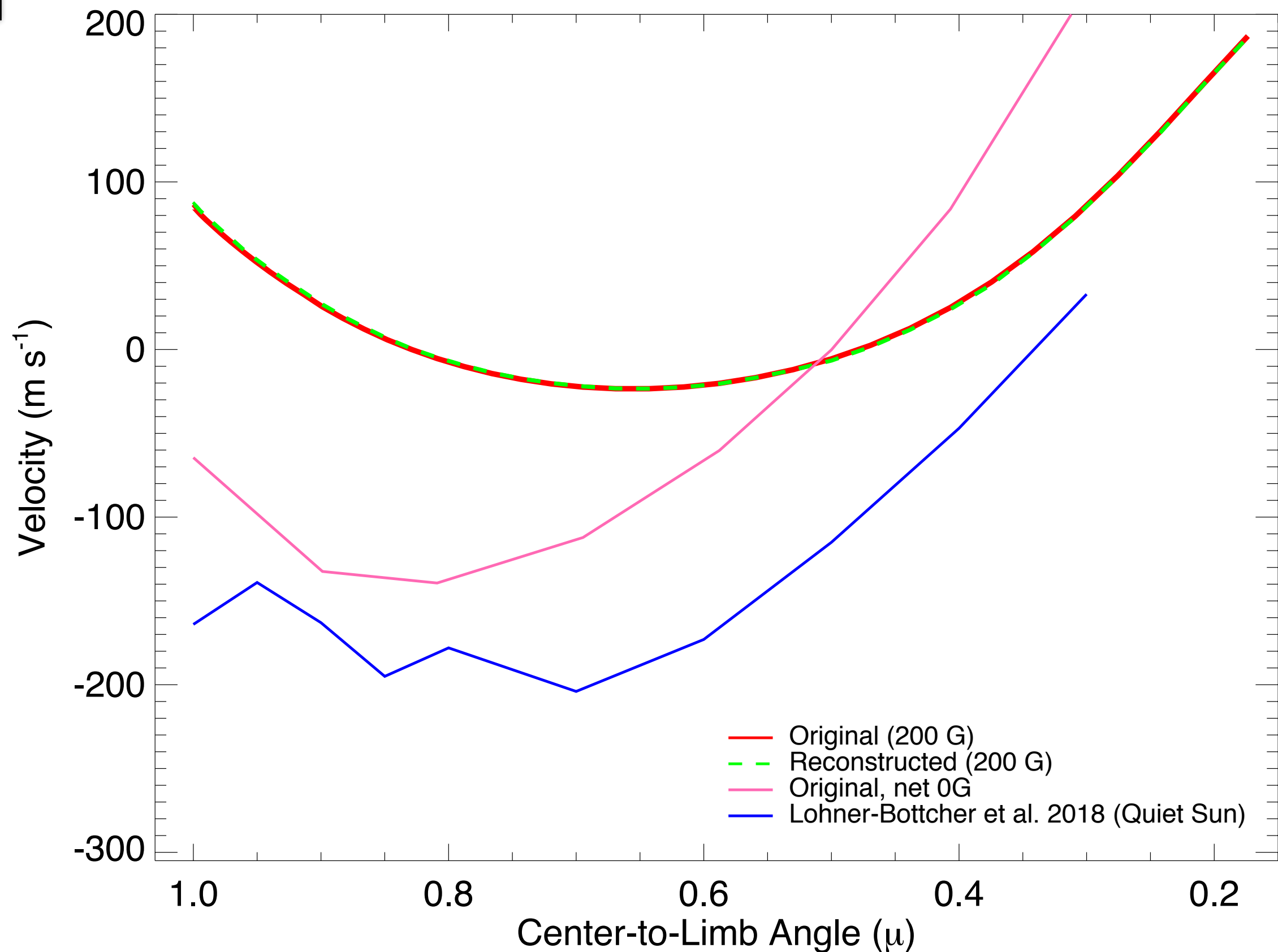


Centre-to-limb Variations on the Sun



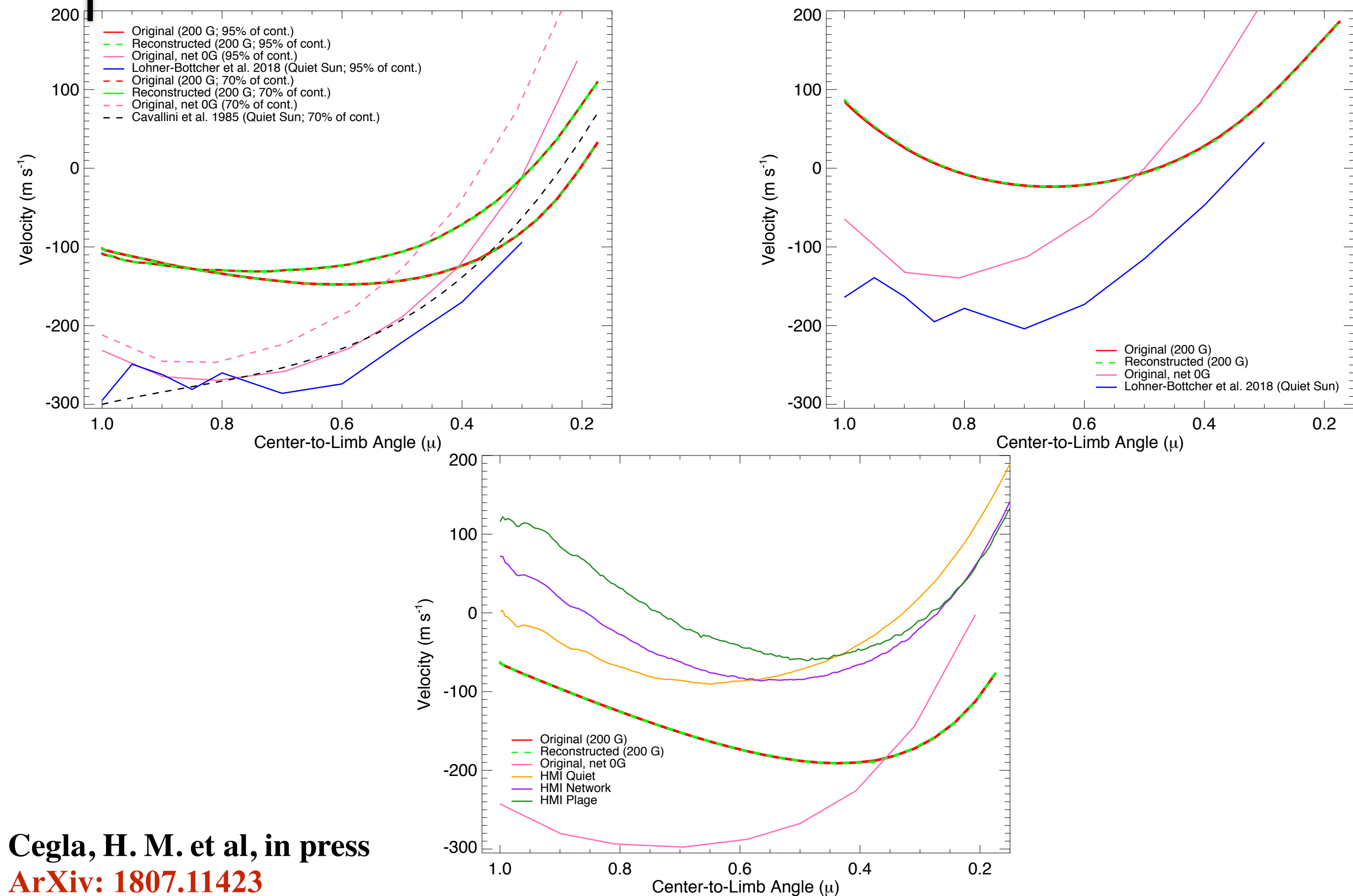


Centre-to-limb Variations on the Sun

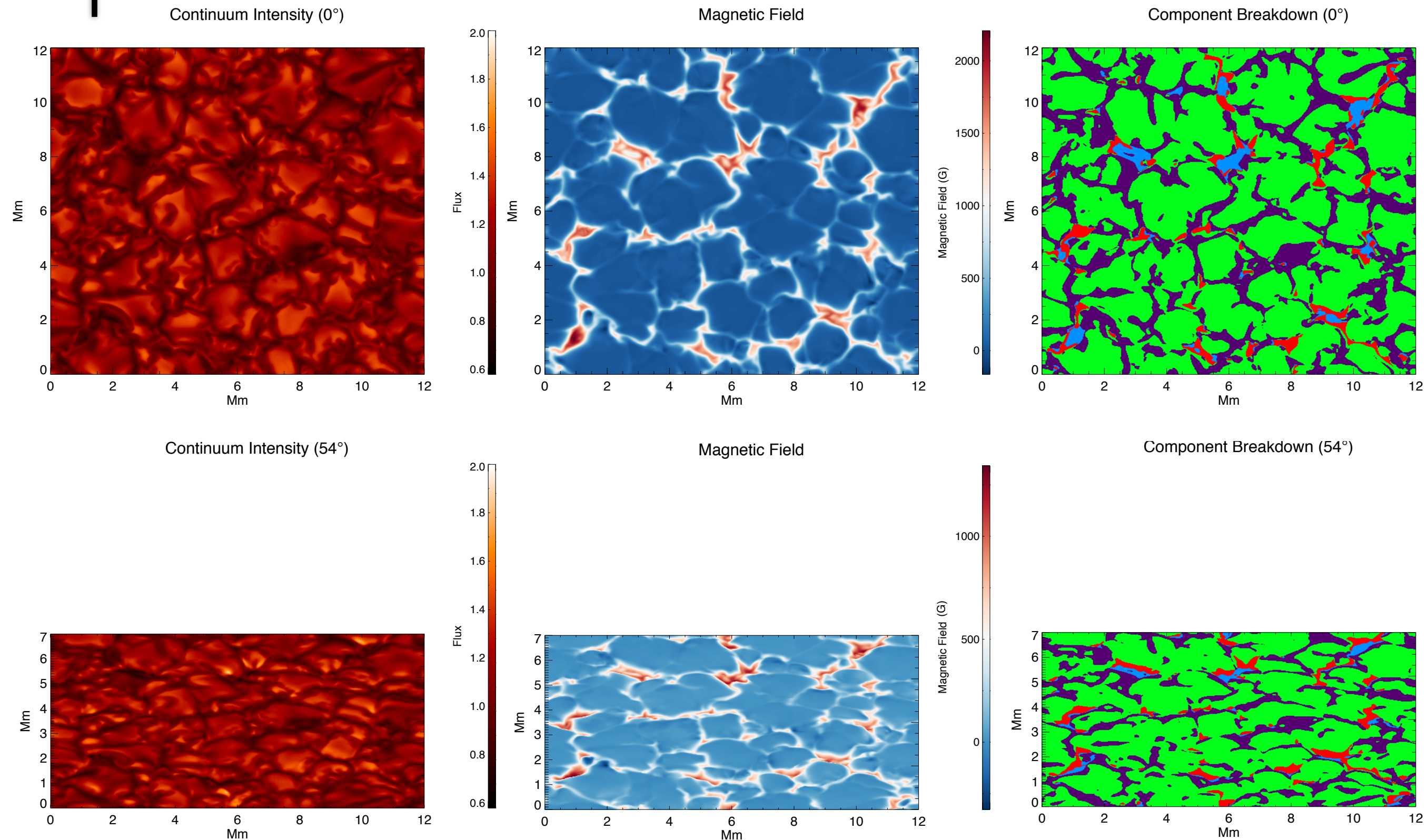




Centre-to-limb Variations on the Sun

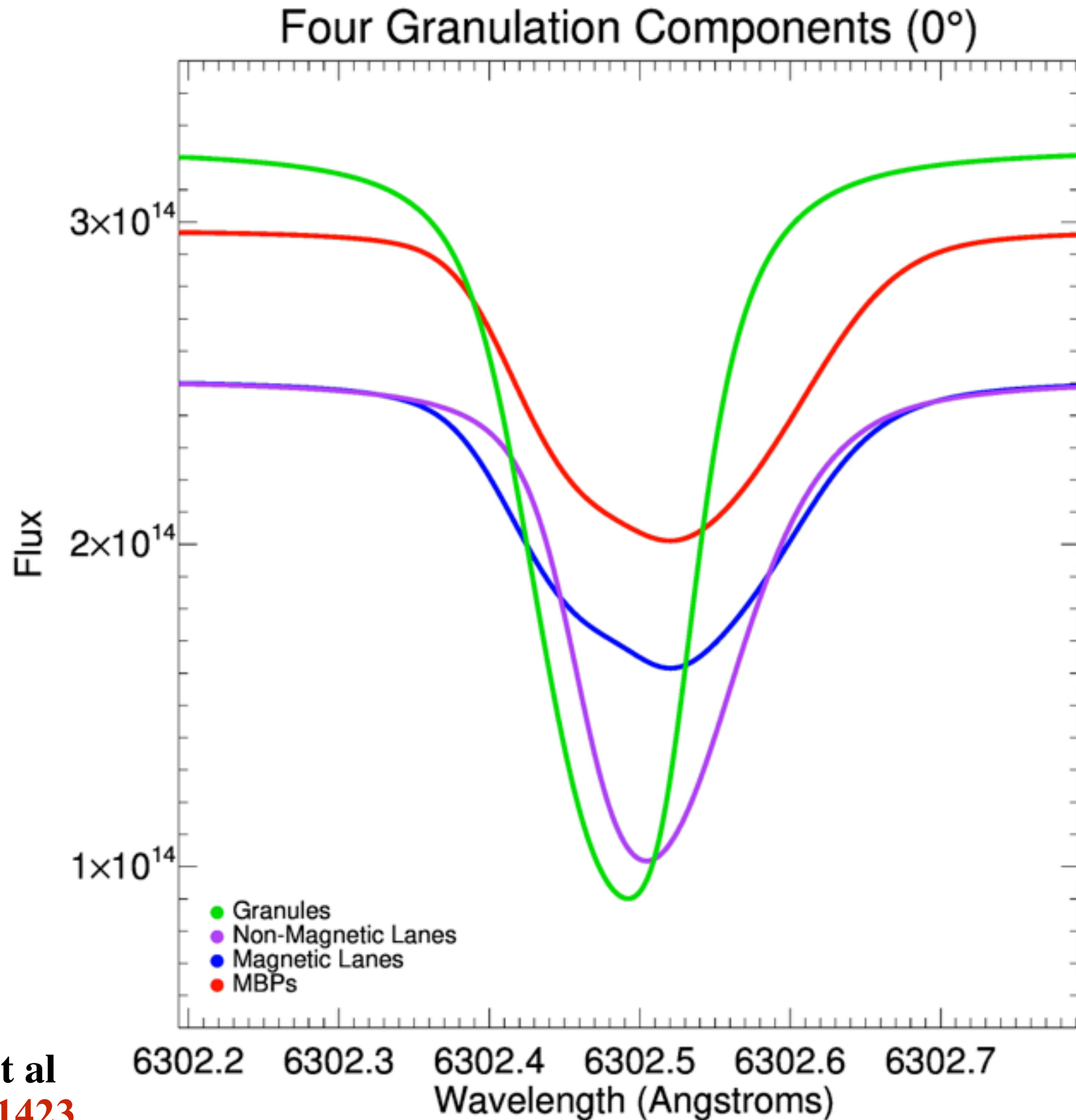


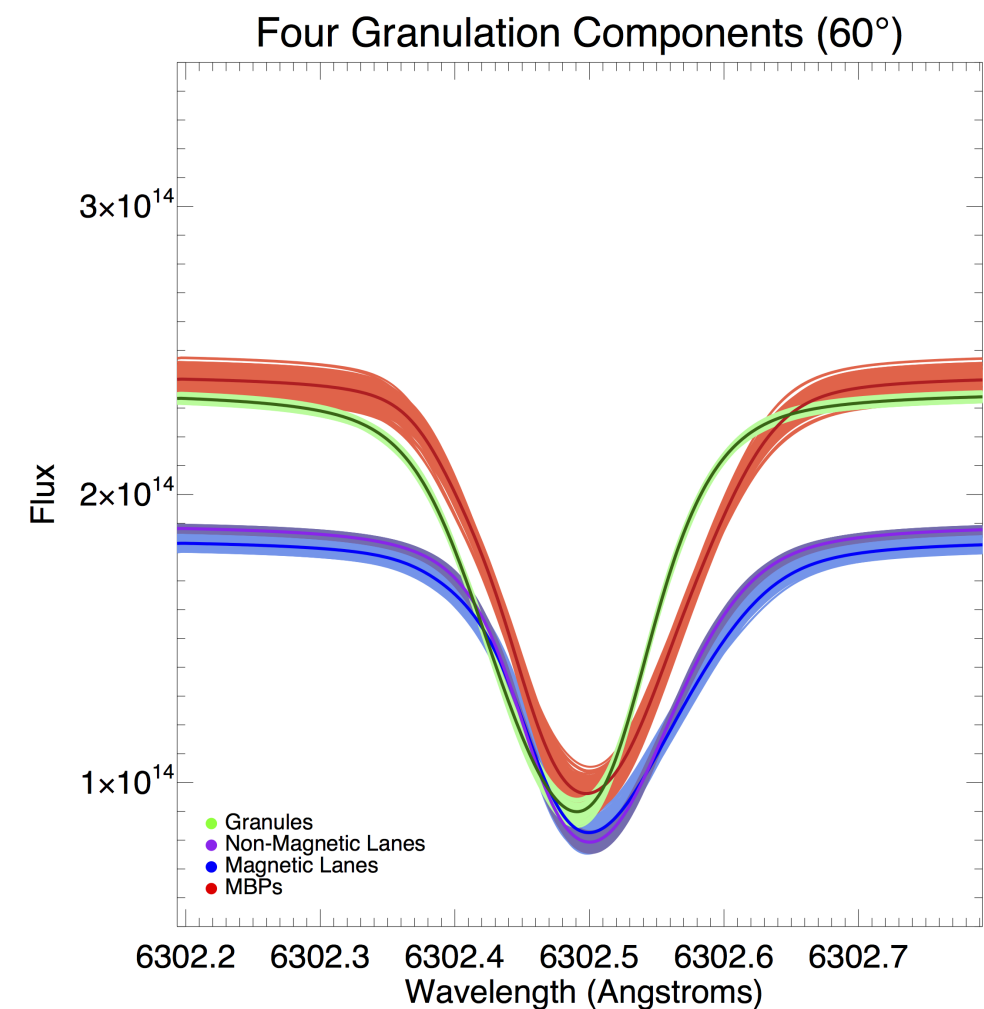
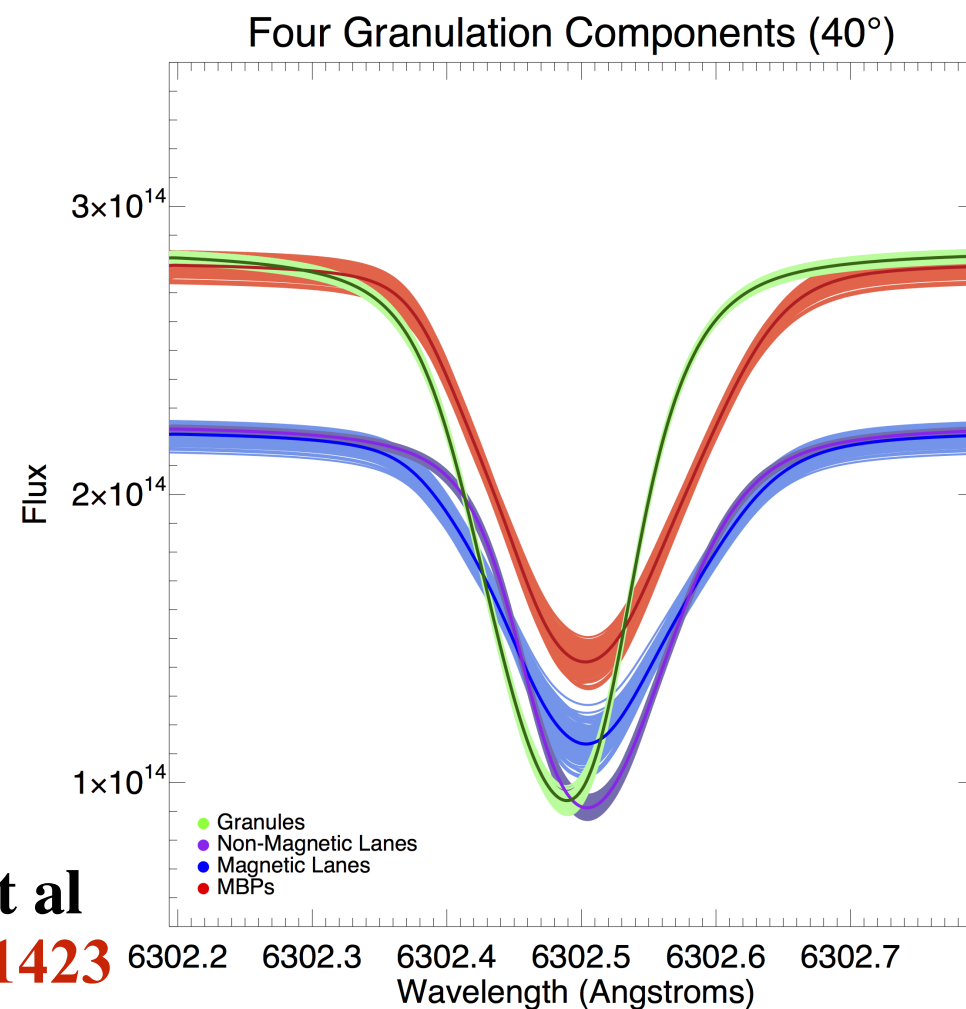
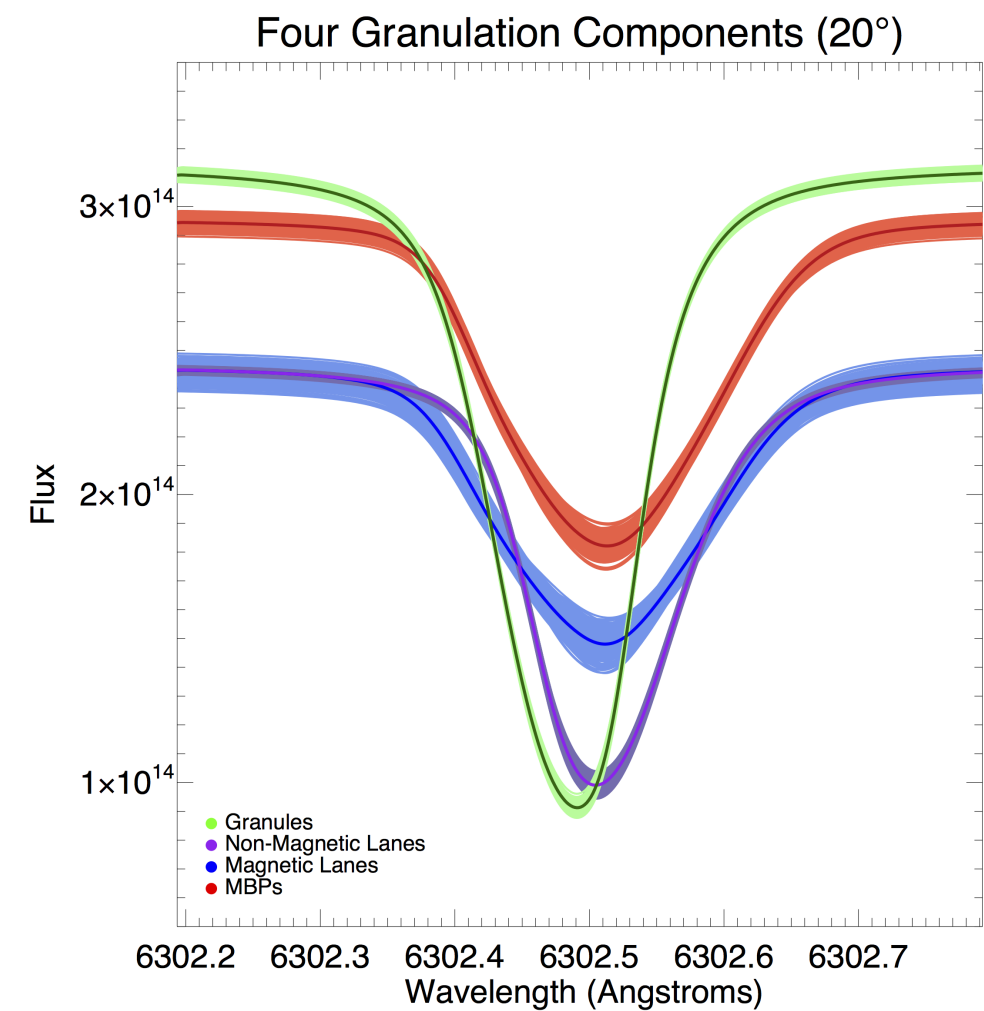
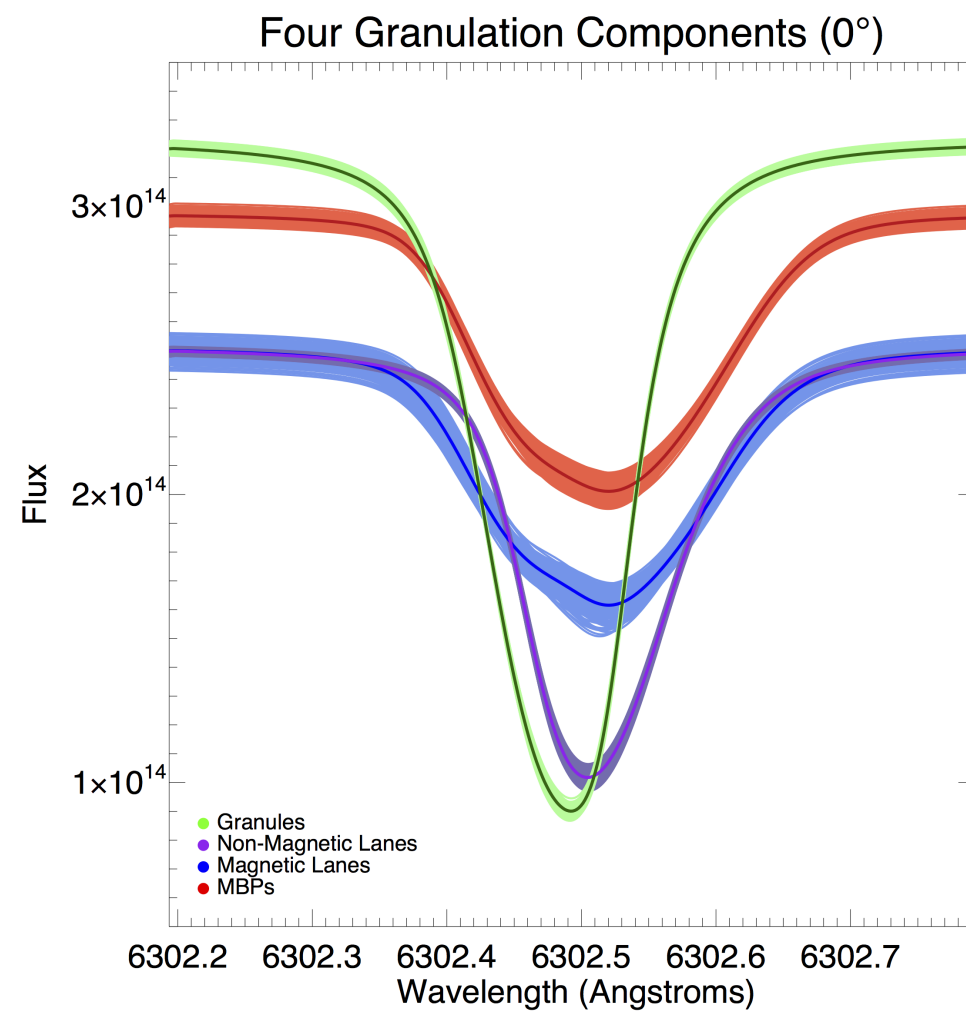
Parameterising the Granulation





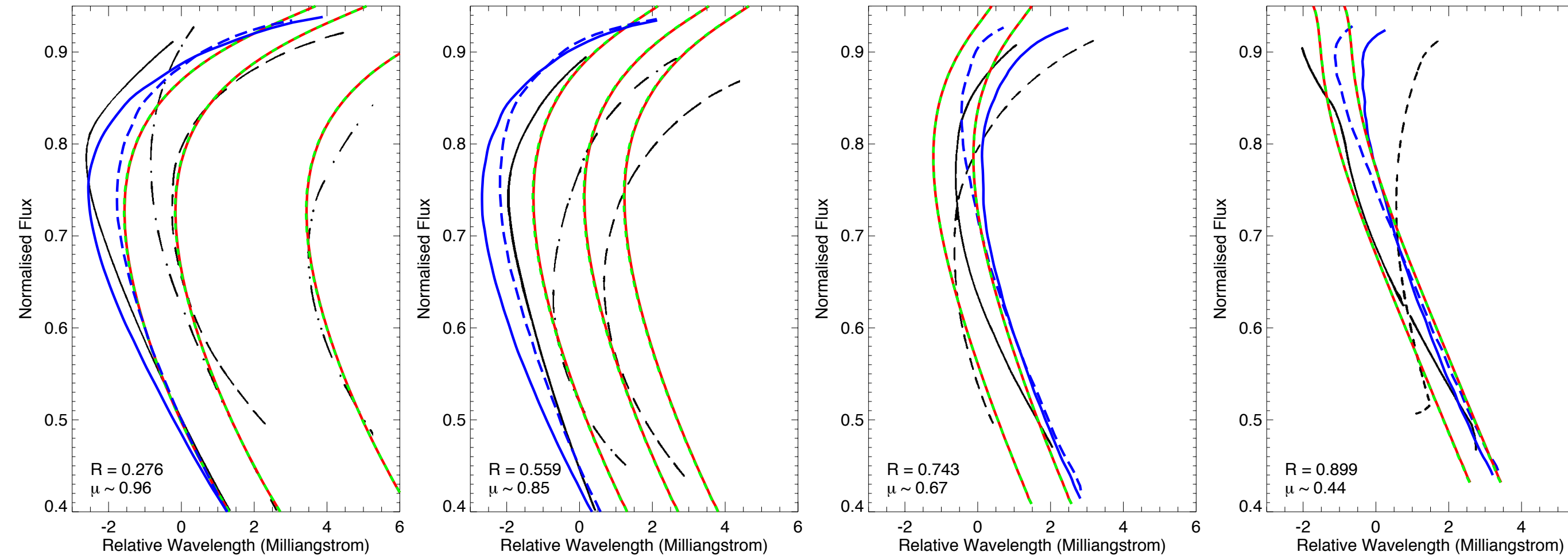
Centre-to-limb Variations on the Sun



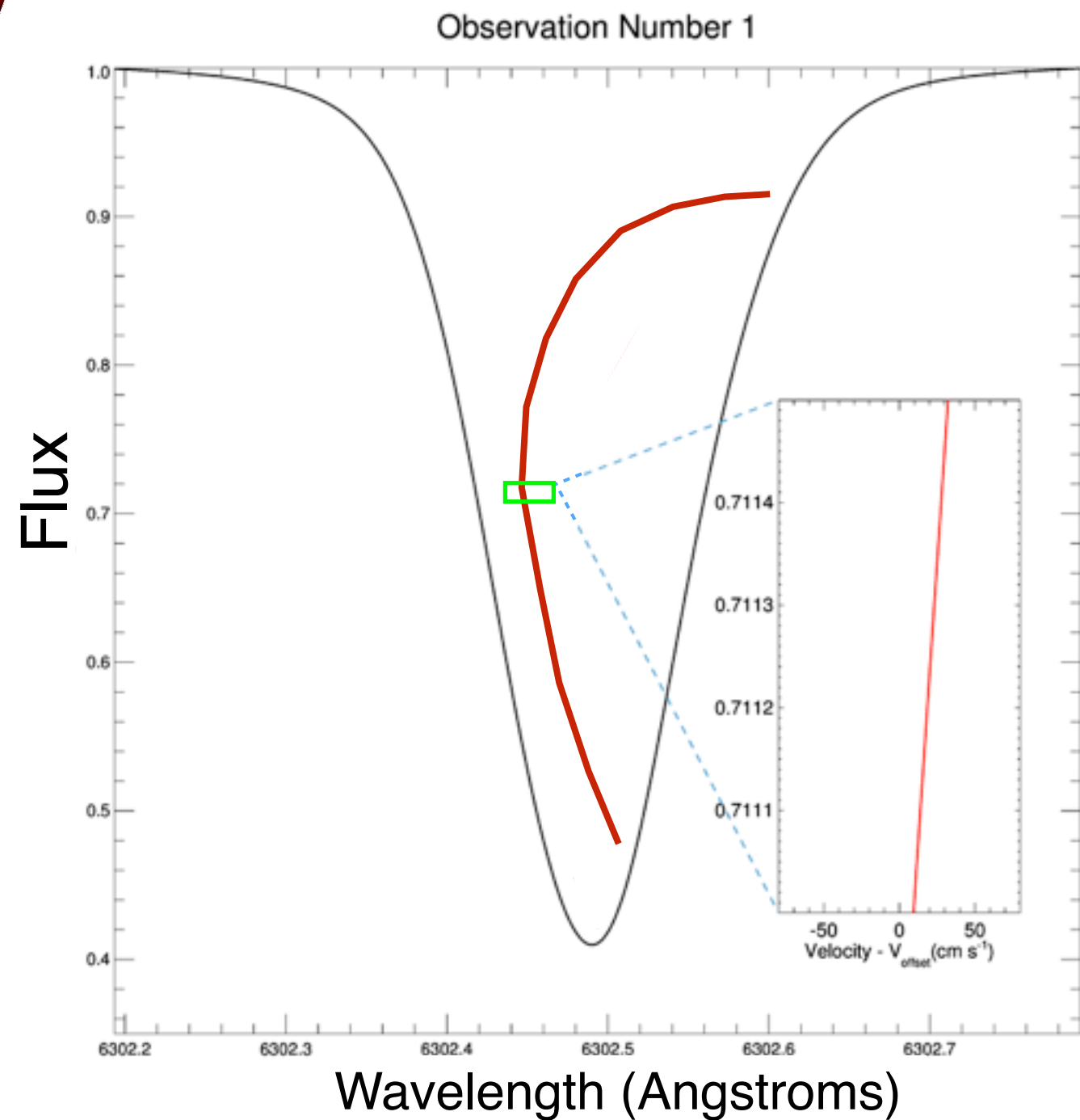
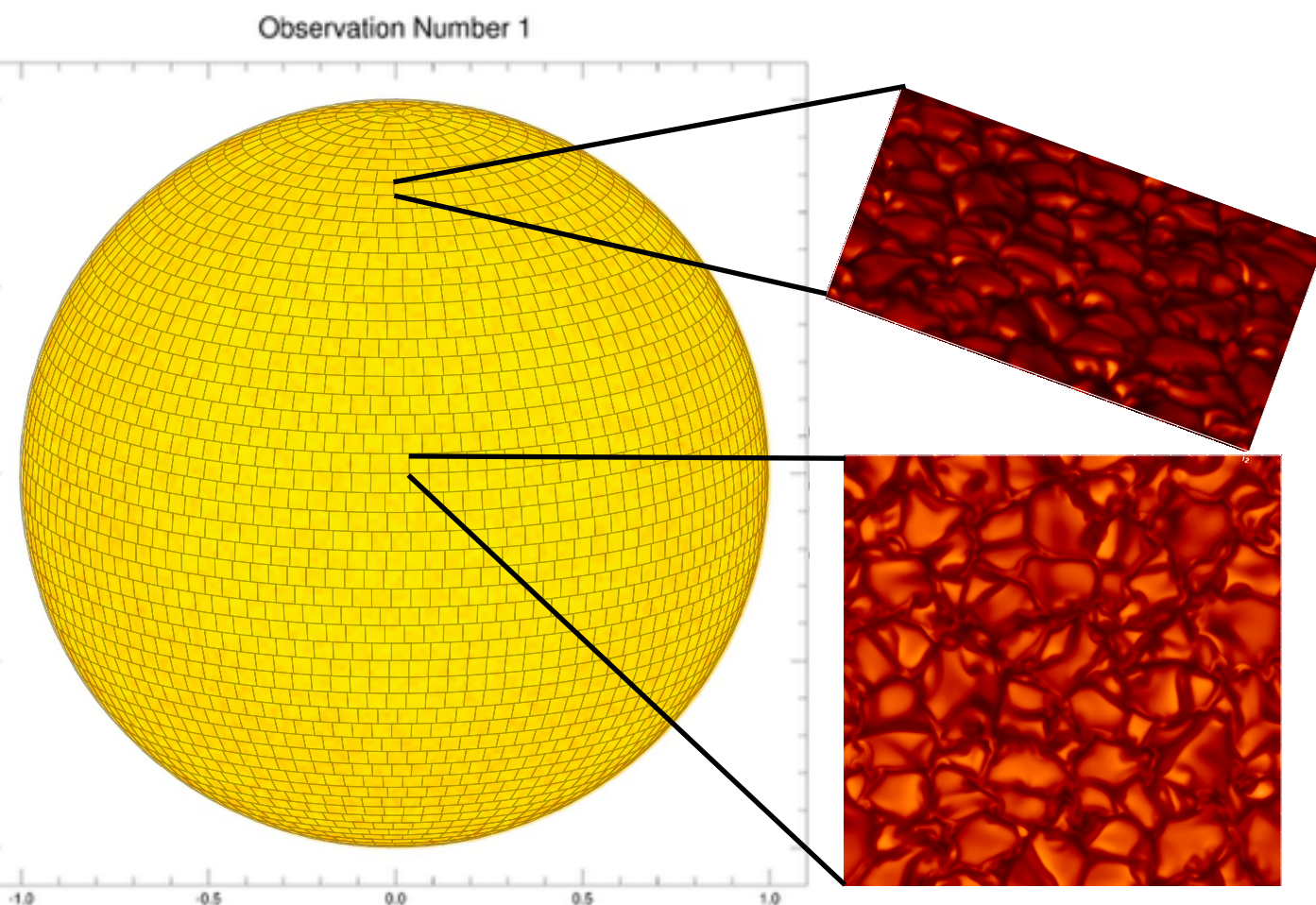




Centre-to-limb Variations on the Sun

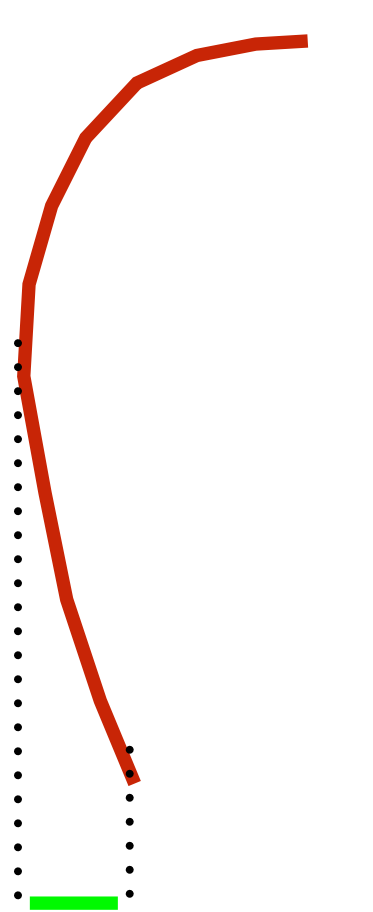
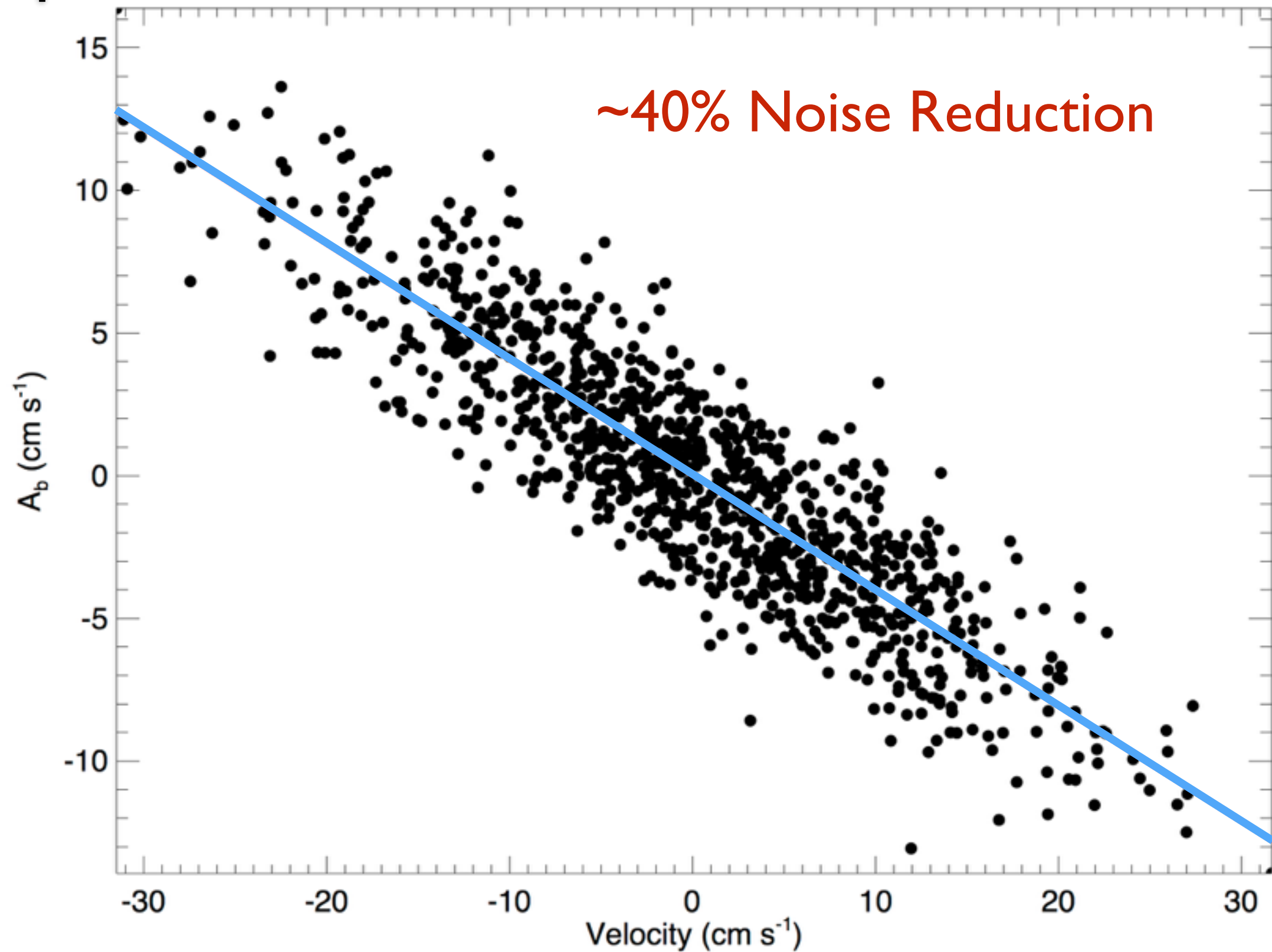


Sun-as-a-star Model Observations



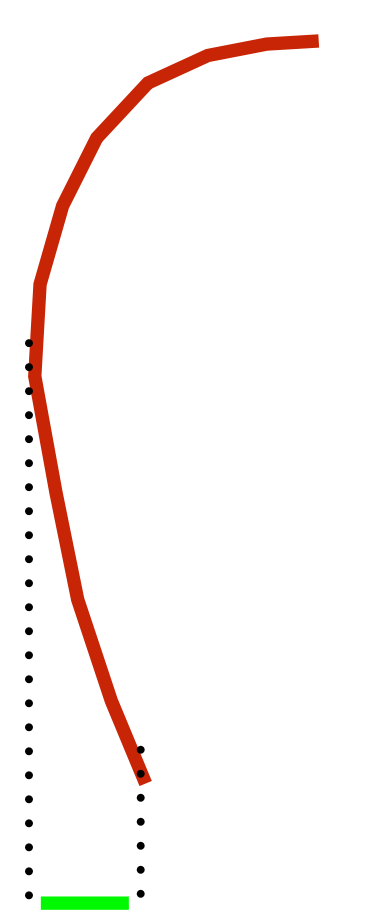
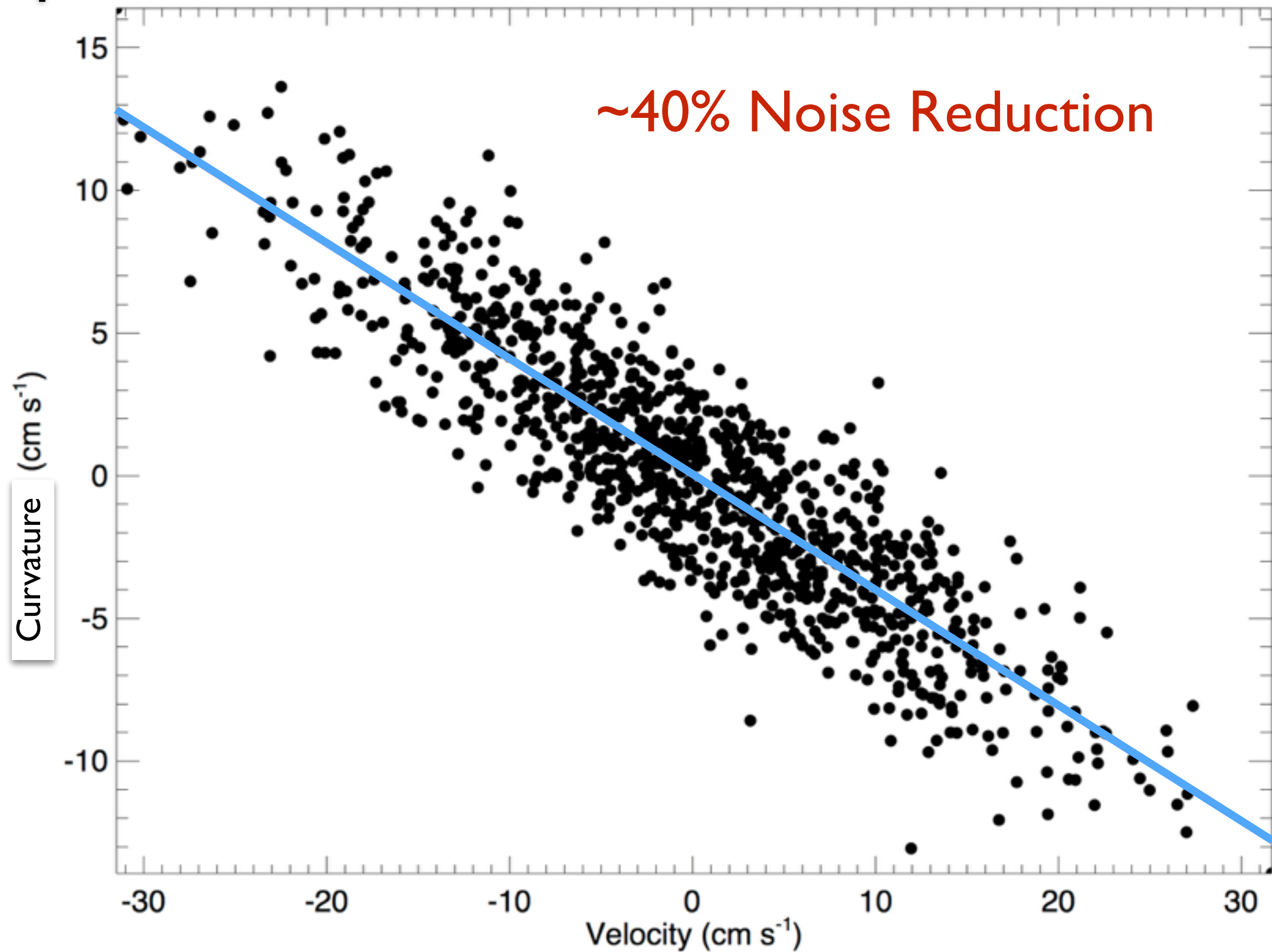


Correlations

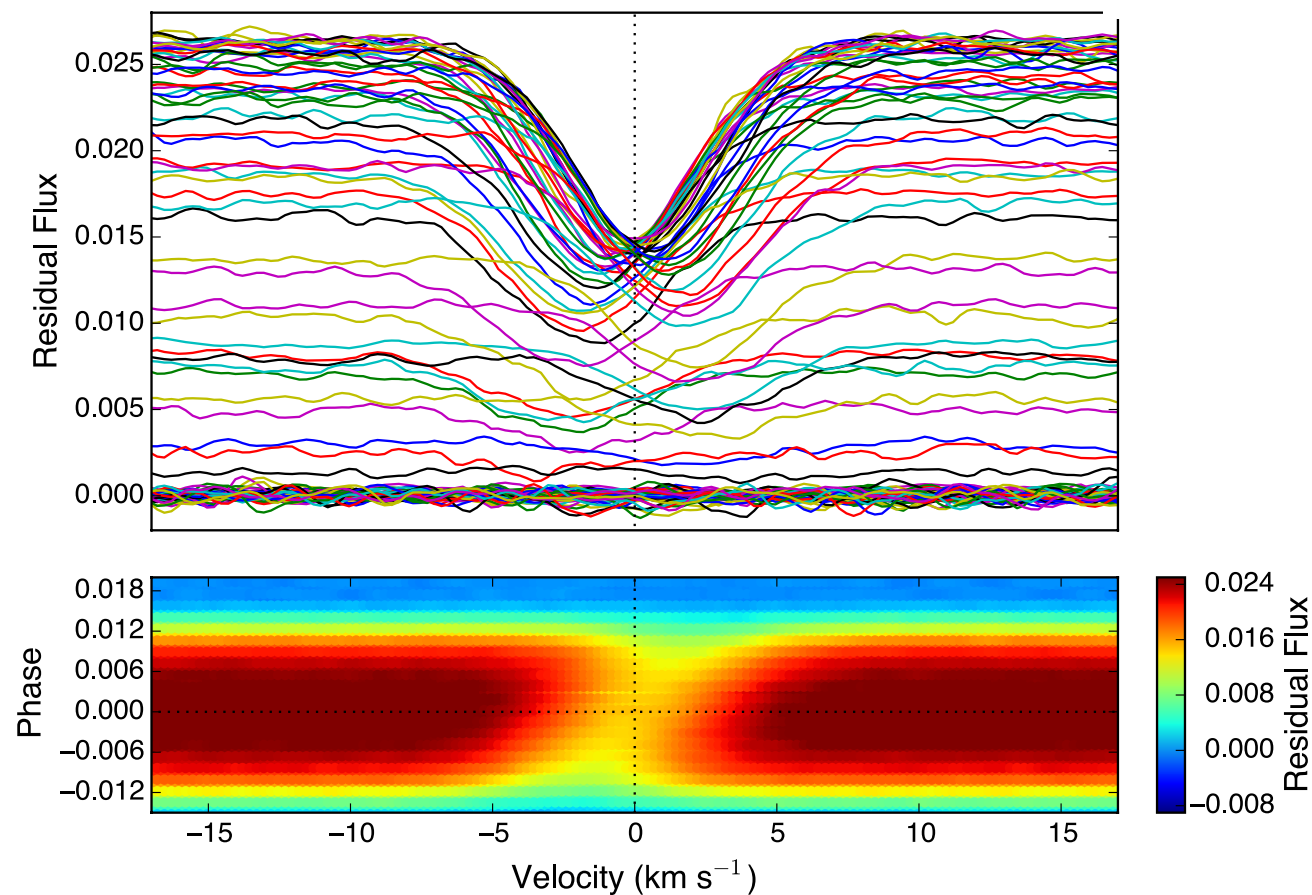
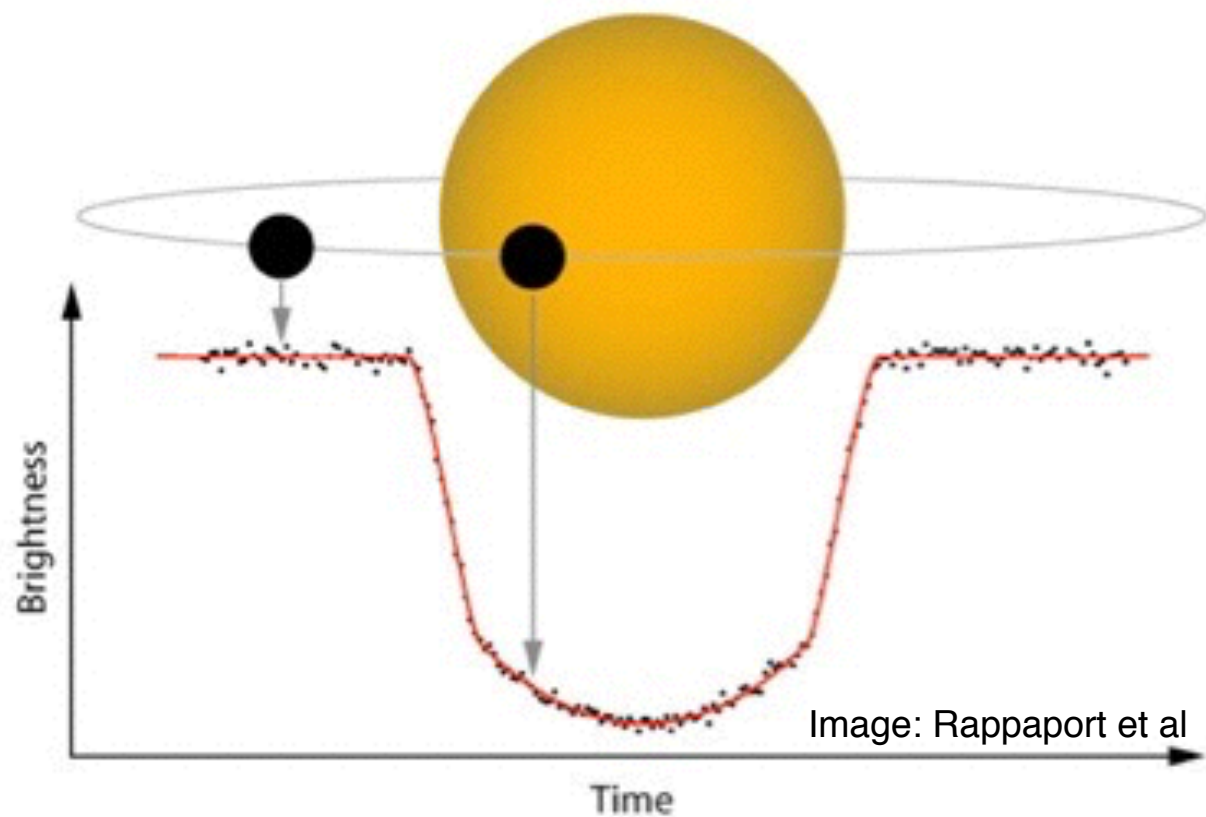
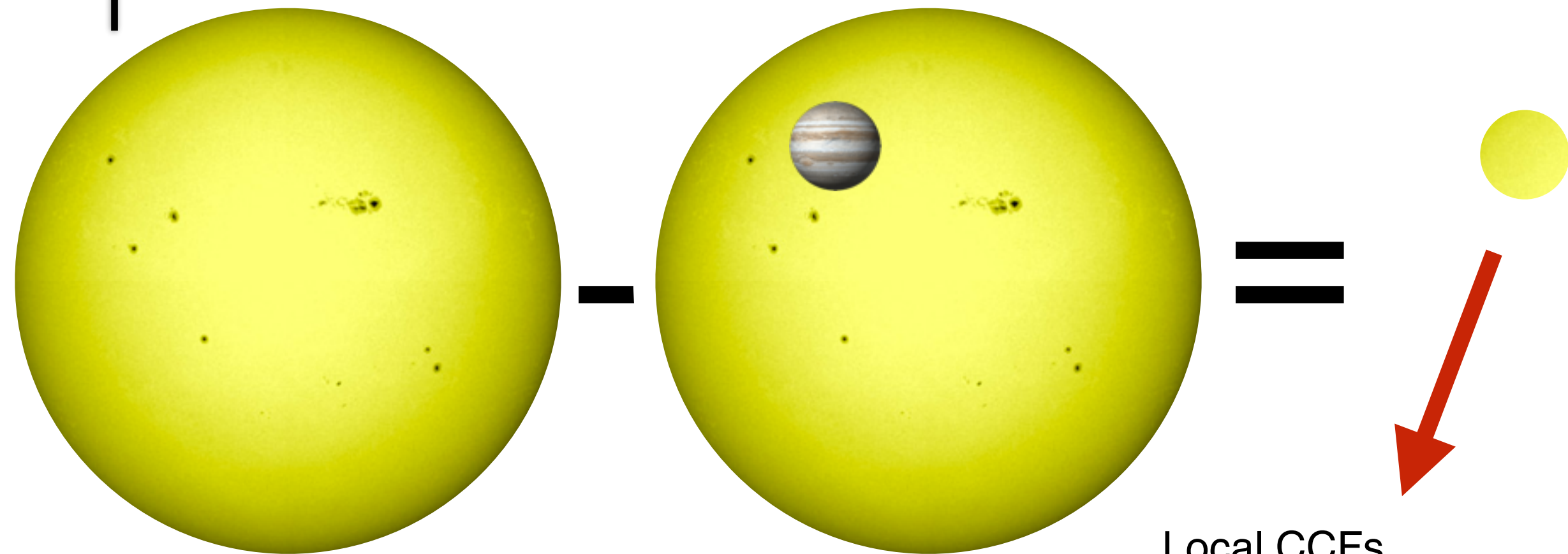




Correlations

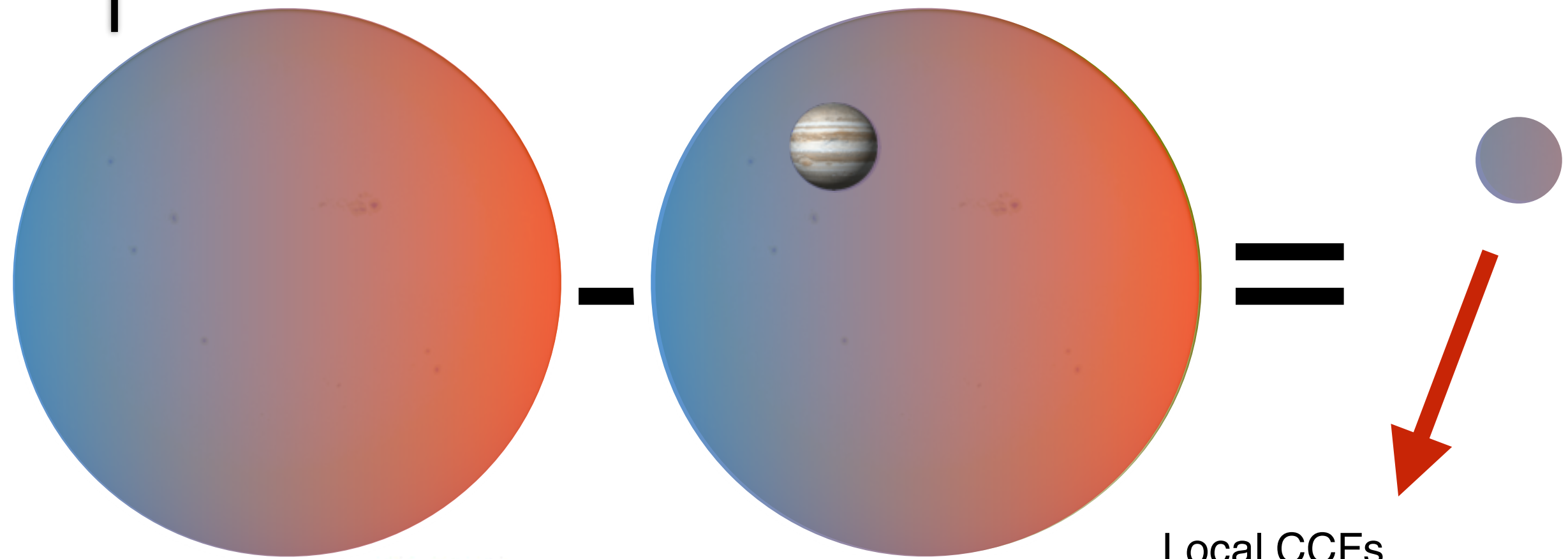


Planets as Probes

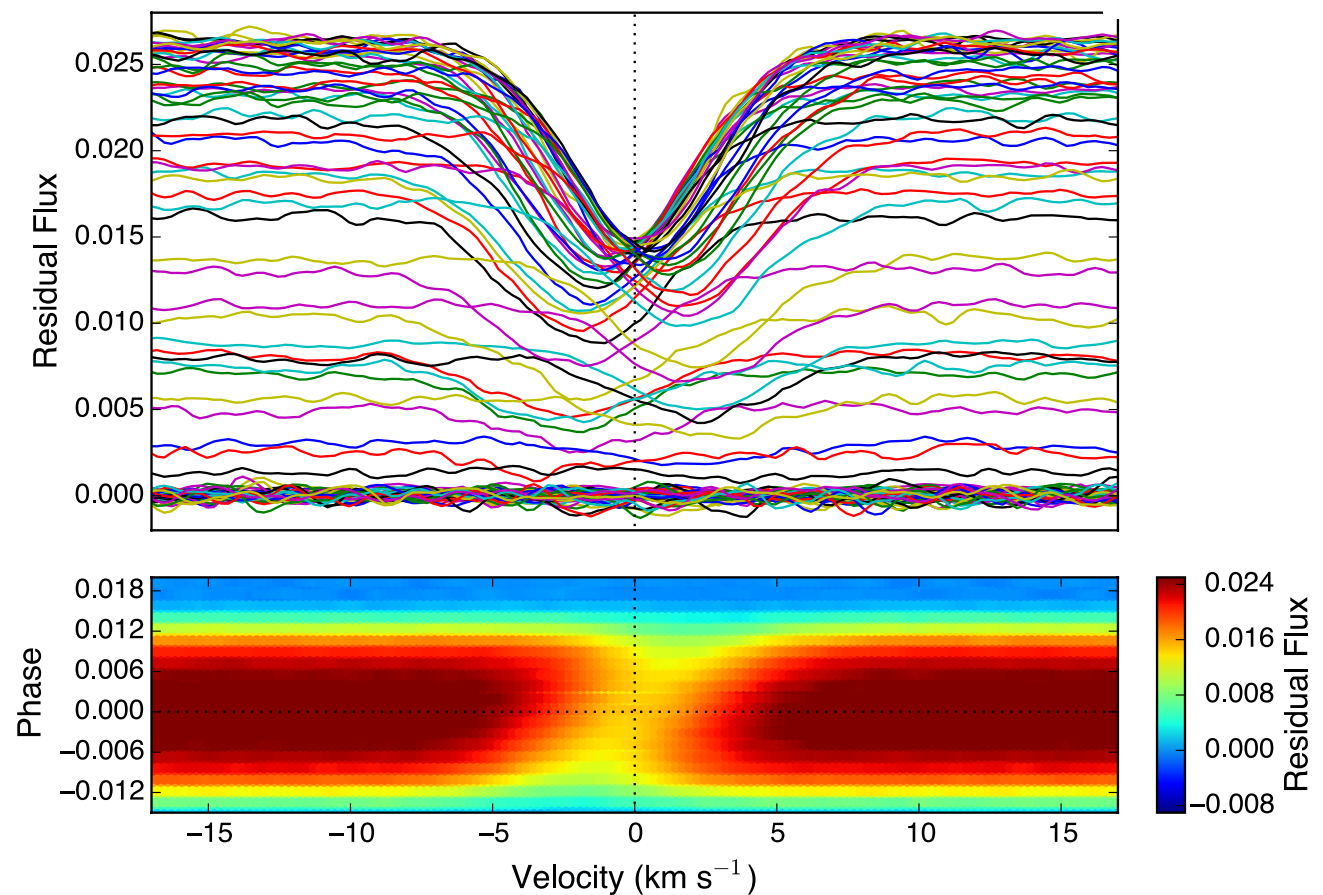
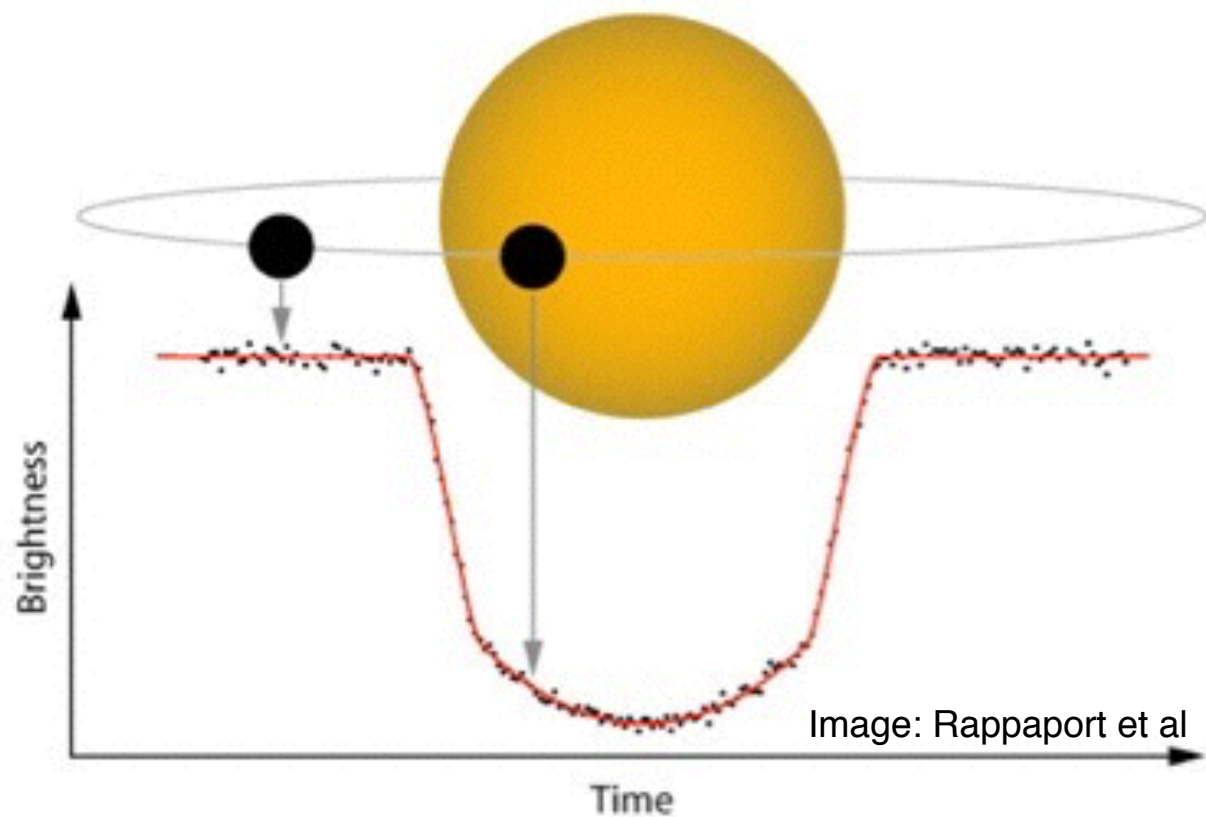


The Rossiter-McLaughlin effect

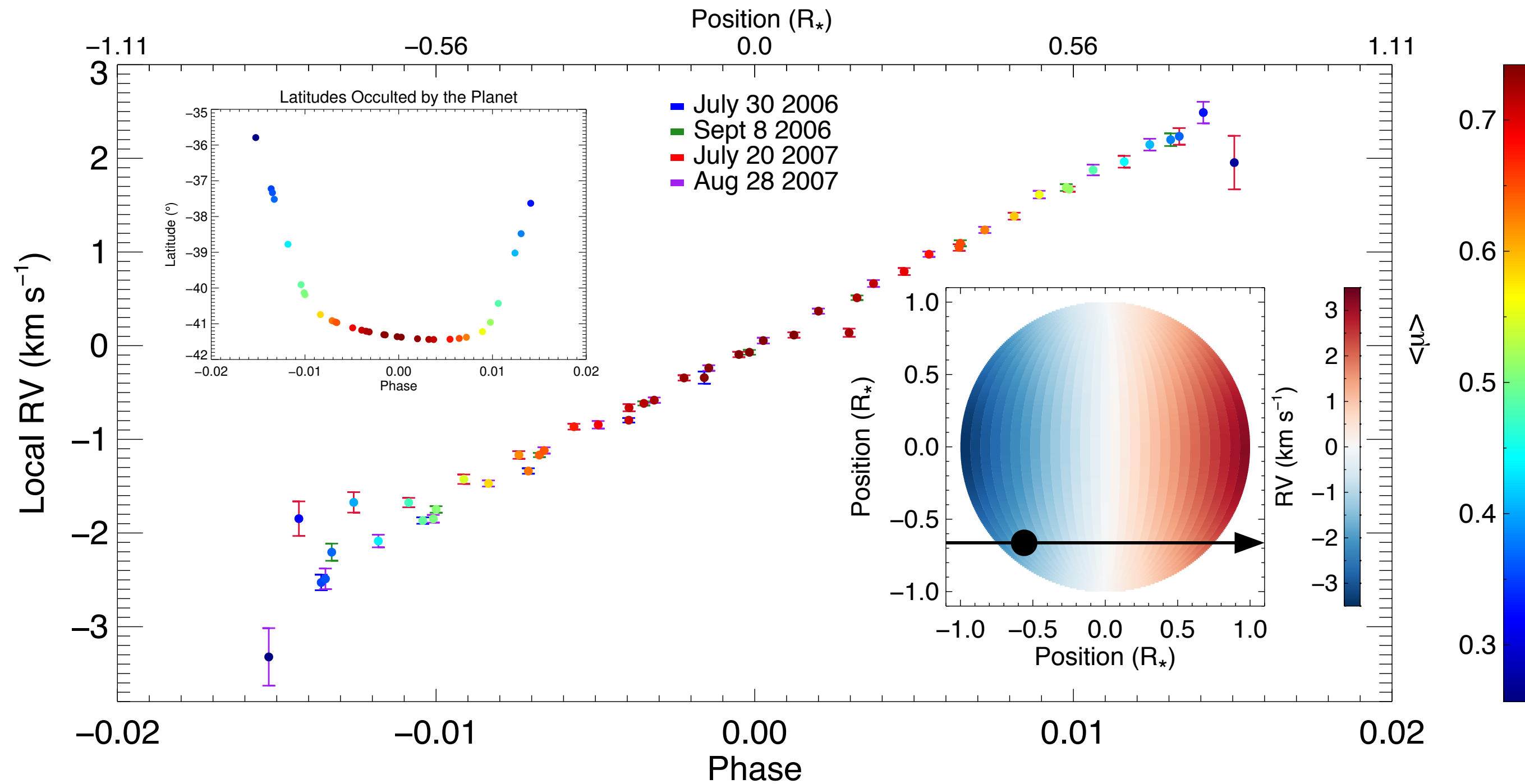
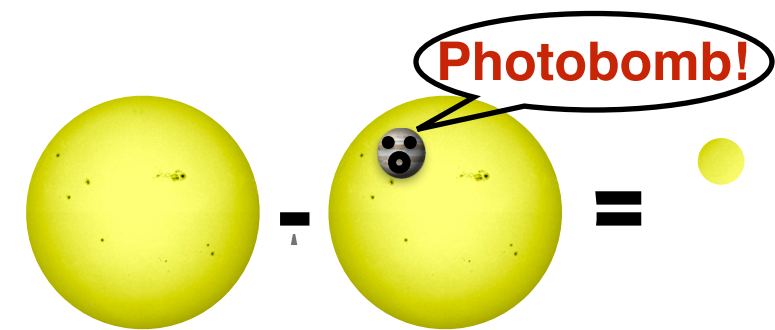
RELOADED



Local CCFs



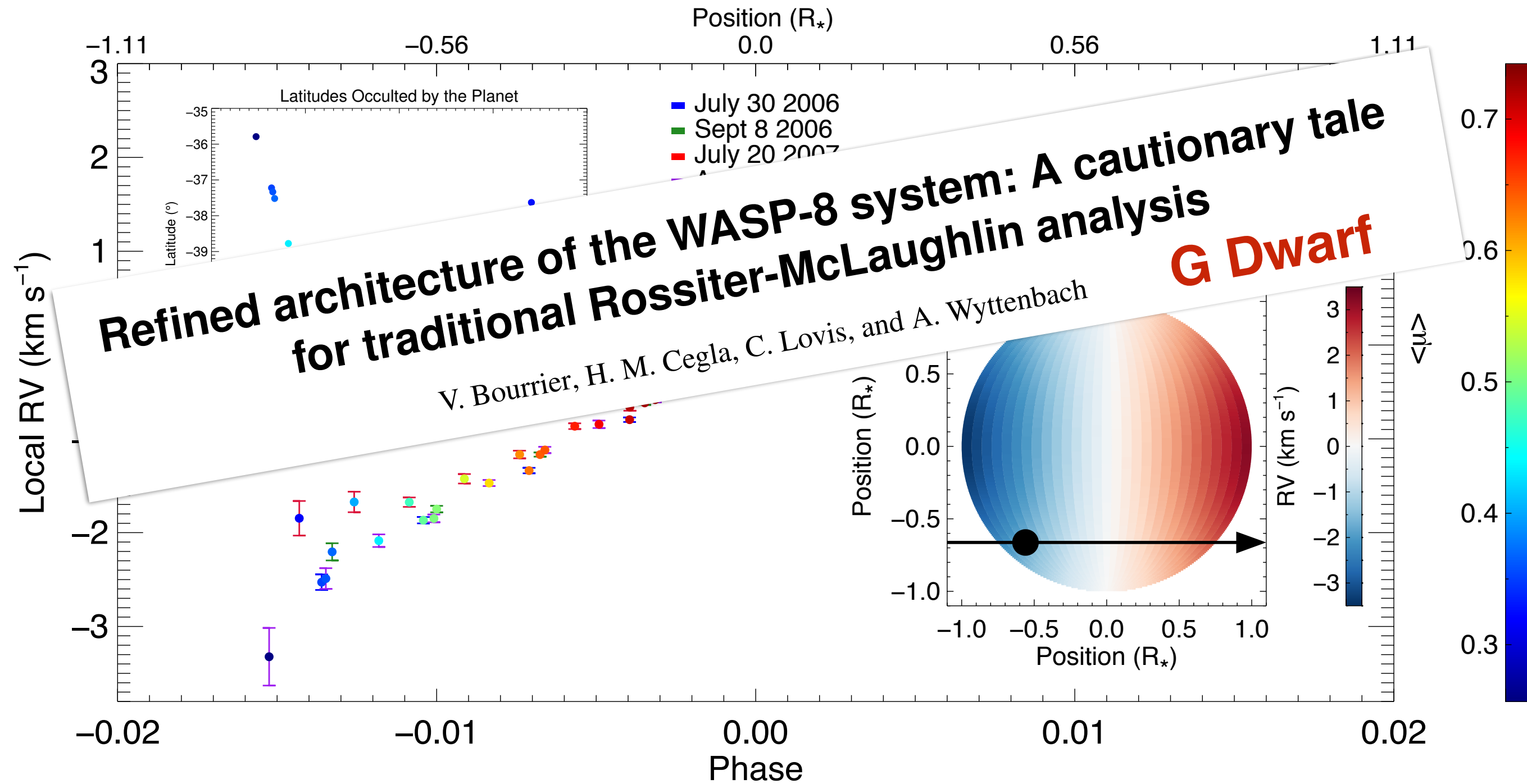
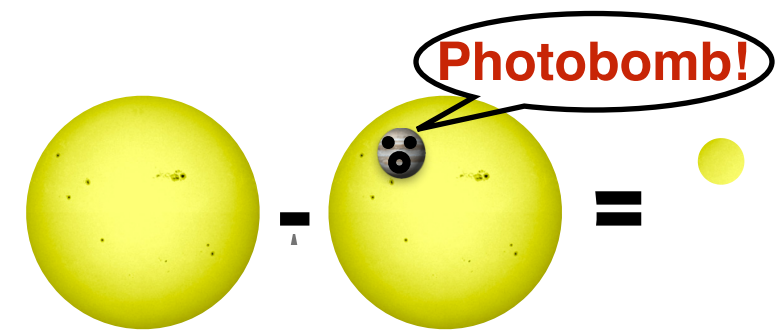
Planets as Probes: HD 189733



$$v_{\text{eq}} \approx 4.5^{+0.5}_{-0.4} \text{ km s}^{-1}; \alpha : 0.3-0.9; > 0.1; i_{\star} \approx 92^{+12}_{-14}^{\circ};$$

$$\lambda \approx -0.4 \pm 0.2^{\circ}; \psi \approx 7^{+12}_{-4}^{\circ}$$

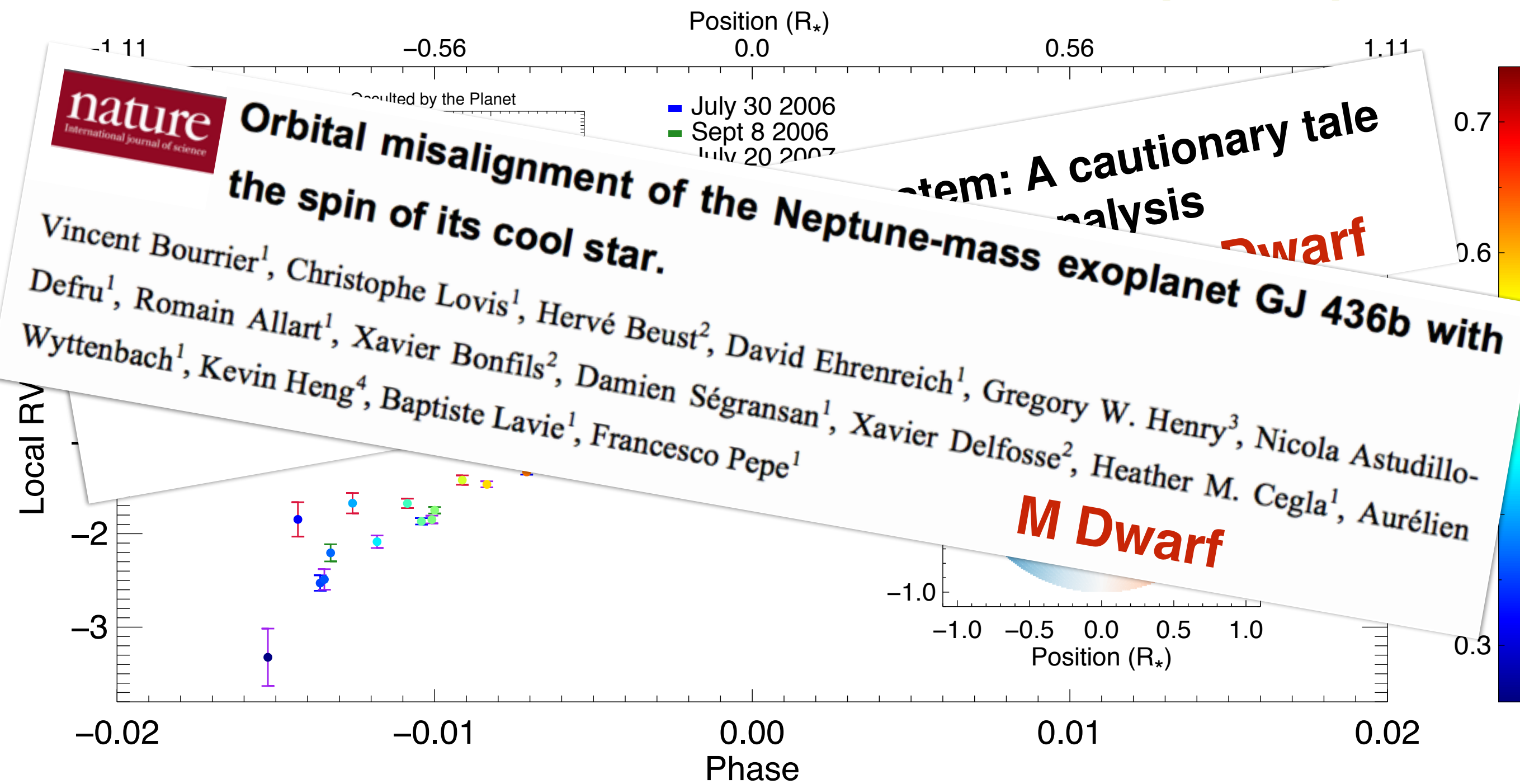
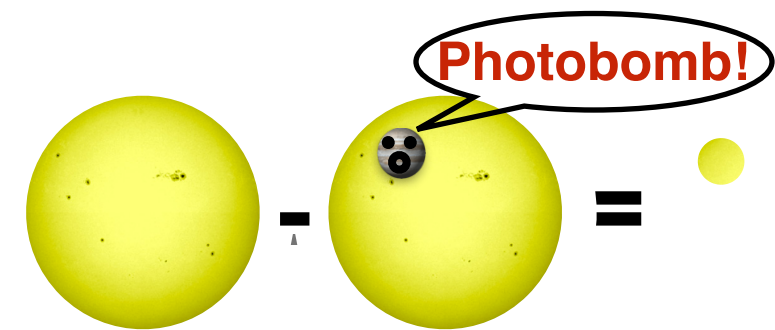
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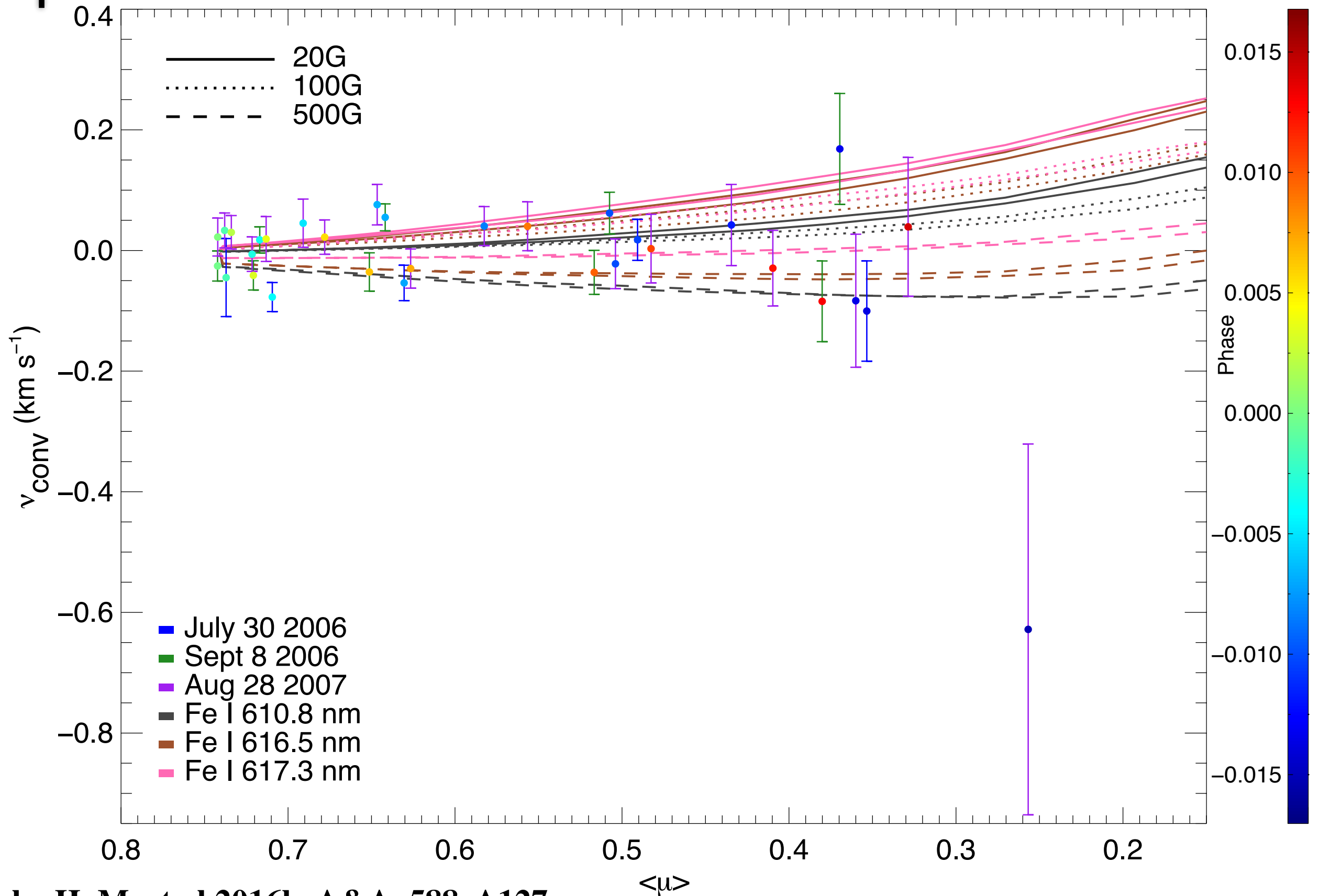
Planets as Probes: HD 189733



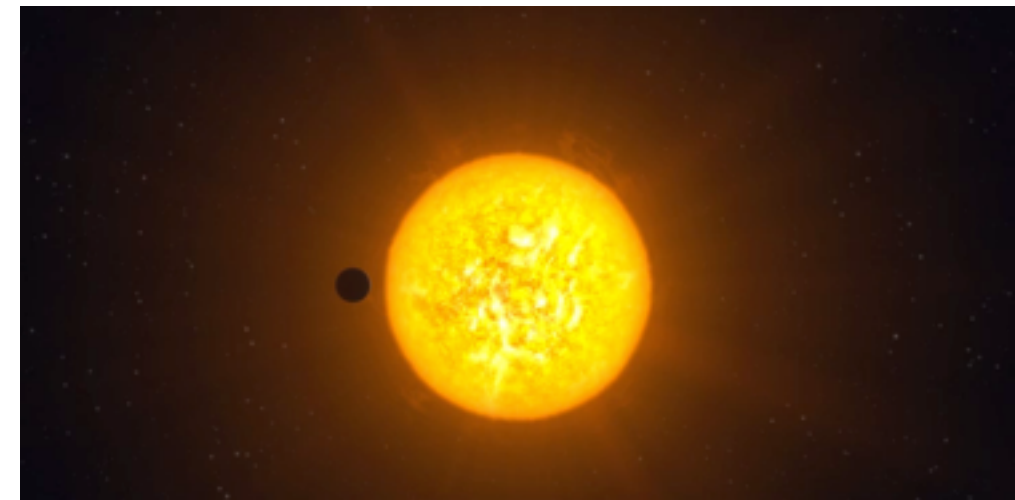
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Planets as Probes: HD 189733



- Stellar surface phenomena alter line profiles & RVs
 - Impacts planet detection/confirmation/characterisation
 - Poses fundamental RV precision limit
 - MHD simulations offer pathway to characterise and disentangle
- Can use planets to spatially resolve stars
 - Probe convection, differential rotation etc.
 - Validate MHD simulations (beyond the Sun!)
 - Study evolution of star-planet systems
- Ask me about oscillations!



What about oscillations...?

